

# Validation of Cross Sections for Monte Carlo Simulation of the Photoelectric Effect

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**Abstract**—Several total and partial photoionization cross section calculations, based on both theoretical and empirical approaches, are quantitatively evaluated with statistical analyses using a large collection of experimental data retrieved from the literature to identify the state of the art for modeling the photoelectric effect in Monte Carlo particle transport. Some of the examined cross section models are available in general purpose Monte Carlo systems, while others have been implemented and subjected to validation tests for the first time to estimate whether they could improve the accuracy of particle transport codes. The validation process identifies Scofield's 1973 non-relativistic calculations, tabulated in the Evaluated Photon Data Library (EPDL), as the one best reproducing experimental measurements of total cross sections. Specialized total cross section models, some of which derive from more recent calculations, do not provide significant improvements. Scofield's non-relativistic calculations are not surpassed regarding the compatibility with experiment of K and L shell photoionization cross sections either, although in a few test cases Ebel's parameterization produces more accurate results close to absorption edges. Modifications to Biggs and Lighthill's parameterization implemented in Geant4 significantly reduce the accuracy of total cross sections at low energies with respect to its original formulation. The scarcity of suitable experimental data hinders a similar extensive analysis for the simulation of the photoelectron angular distribution, which is limited to a qualitative appraisal.

**Index Terms**—Geant4, Monte Carlo, simulation, x-rays.

## I. INTRODUCTION

PHOTOIONIZATION is important in various experimental domains, such as material analysis applications, astrophysics, photon science and bio-medical physics. As one of the interactions photons undergo in matter, it is relevant in experimental methods concerned with the energy deposition resulting from photons as primary or secondary particles. Photoionization is also experimentally relevant for the secondary atomic processes that it induces, X-ray

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fluorescence and Auger electron emission, which play a role in many physics research areas and technological applications. Extensive reviews, which cover both the theoretical and experimental aspects of this process, can be found in the literature: [1]–[8] are some notable examples among them.

This paper is concerned with modeling the physics of photoionization in a pragmatic way: the simulation of this process in general purpose Monte Carlo codes for particle transport.

Calculations for the simulation of the photoelectric effect are implemented in these codes [9], nevertheless a comprehensive, quantitative appraisal of their validity is not yet documented in the literature. Assessments previously reported in the literature concern comparisons of cross sections with NIST reference values, such as [10], [11], or comparisons of cross section libraries used in Monte Carlo codes [12], or involve complex observables resulting from several physics processes in the full simulation of an experimental set-up, such as [13]. Comparisons with experimental data of basic modeling features of photoionization in Monte Carlo codes, such as those shown in [14], are usually limited to qualitative visual appraisal and to a restricted sample of photon energies and target materials. It is worthwhile to note that the validation of simulation models implies their comparison with experimental measurements [15]; comparisons with tabulations of theoretical calculations or analytical parameterizations, such as those reported in [16] as validation of Geant4 photon interaction cross sections, do not constitute a validation of the simulation software.

The analysis documented here evaluates the methods adopted in widely known Monte Carlo systems for the calculation of photoelectric cross sections for the elements of the periodic table, as well as other modeling approaches not yet implemented in these codes. This investigation aims to assess the capabilities of Monte Carlo codes for particle transport in this respect, and identify the state of the art for the simulation of the photoelectric effect.

Special emphasis is devoted to the validation and possible improvement of photoionization simulation in Geant4 [17], [18]; nevertheless, the results documented in this paper provide information relevant to other Monte Carlo systems as well.

The simulation of the atomic relaxation following the ionization of an atom has been treated in previous publications [19]–[22], therefore it is not addressed in this paper.

## II. STRATEGY OF THIS STUDY

This study concerns an extensive set of models for the simulation of photoionization, which are representative of the variety

of theoretical and empirical methods documented in the literature. The validation test concerns single ionisation of neutral atoms by non-polarized photons, as this is the context handled by general purpose Monte Carlo codes.

Computational performance imposes constraints on the complexity of physics calculations to be performed in the course of simulation; hence the analysis in this paper is limited to theoretical cross sections for which tabulations of pre-calculated values are available and to empirical models that are expressed by means of simple analytical formulations. To be relevant for general purpose Monte Carlo systems, tabulated data should cover the whole periodic table of elements and an extended energy range.

The evaluation mainly concerns total and partial photoelectric cross sections: in particle transport, the former are relevant to determine the occurrence of the photoionization process, while the latter determine which shell is ionized. Calculated cross sections are quantitatively compared with a wide set of experimental data collected from the literature. The compatibility with experiment for each model, and the differences in compatibility with experiment across the various models, are quantified by means of statistical methods.

In addition, methods for the determination of the photoelectron angular distribution are examined; nevertheless, due to the scarcity of pertinent experimental data, their analysis is limited to qualitative considerations.

Computational algorithms pertaining to how basic physics modeling features are used in the transport environment, such as methods for dealing with the macroscopic cross sections for compounds or mixtures [23], [24], are outside the scope of this paper.

### III. PHYSICS OVERVIEW

Photoionization has been the object of theoretical and experimental interest for several decades; only a brief overview of the physics relevant to the simulation of the photoelectric effect in general purpose Monte Carlo codes is included here to facilitate the understanding of the validation tests reported in this paper.

In the photoelectric effect a photon disappears and an electron is ejected from an atom. The energy of the photoelectron corresponds to the difference between the energy of the absorbed photon and the energy binding the electron to the atom.

General purpose Monte Carlo codes for particle transport consider single photon interactions with isolated atoms in their ground state; they neglect interactions with ions and excited states, and multiple ionizations. Photon interactions are treated regardless of the environment of the target medium; this assumption neglects solid state effects and other features related to the molecular structure of the medium. The environment can have a significant effect on the cross sections near the photoionization thresholds of both inner and outer shell electrons; due to the limitations of their underlying physics assumptions, current general purpose Monte Carlo codes are not usually exploited for the simulation of experimental scenarios involving EXAFS (Extended X-ray Absorption Fine Structure), XANES

(X-ray Absorption Near Edge Structure) and other techniques for which detailed accounting of material structure is required.

#### A. Total and Partial Cross Sections

The photoelectric cross section as a function of energy exhibits a characteristic sawtooth behavior corresponding to absorption edges, as the binding energy of each electron subshell is attained and corresponding photoionization is allowed to occur.

Early theoretical calculations of photoionization cross sections were limited to the K shell; they are represented by the papers of Pratt [25], providing the asymptotic behavior for arbitrarily high energies, and Pratt et al. [26], reporting calculations in the energy range between 200 keV and 2 MeV. More extensive calculations became available only at a later stage: Rakavy and Ron [27] calculated cross sections for all subshells of five elements over the energy range 1 keV to 2 MeV, and Schmickley and Pratt [28] reported cross sections for K to M shells for three elements from 412 to 1332 keV.

Scofield's non-relativistic calculations [29] in a Hartree-Slater framework represented a major advancement in the field, as they covered systematically all subshells over the whole periodic table of the elements. More recent calculations were performed by Chantler [30], [31] in a self-consistent relativistic Dirac-Hartree-Fock framework.

These theoretical calculations provide the basis for the tabulated data libraries listed in Section IV.

Various empirical formulations of photoionization cross sections are reported in the literature, e.g. in [32]. They derive from fits to experimental data, parameterizations of theoretical calculations and semi-empirical methods involving both measured data and theoretical considerations.

#### B. Photoelectron Angular Distribution

Fischer's non-relativistic theory [33] addresses the calculation of differential cross sections in the low energy region. The first relativistic treatment of the photoelectric effect was given by Sauter [34], [35], who calculated the K-shell cross section in the Born approximation; it concerns the lowest order in  $Z\alpha/\beta$  (where  $Z$  is the atomic number of the target,  $\alpha$  is the fine structure constant and  $\beta$  is  $v/c$ ). A comparison of these theories is discussed in [36], which showed that Sauter's theory applies even in the non-relativistic realm, despite being derived for relativistic electrons.

Gavrila [37] and Nagel [38] extended Sauter's results to the next order in  $Z\alpha/\beta$ . Further calculations by Gavrila are available for the L shell [39].

Monte Carlo codes generate the photoelectron angular distribution based on these differential cross section calculations; their respective approaches are documented in Section V.

### IV. PHOTOIONIZATION CROSS SECTION COMPILATIONS

The photoelectric cross section compilations considered in this study are listed in Table I, which reports the label by which

TABLE I  
COMPILATIONS OF PHOTOIONIZATION CROSS SECTIONS CONSIDERED  
IN THIS STUDY

Identifier	Energy range		Z range		Shell
Biggs [50]	10 eV	100 GeV	1	100	no
BiggsG4 [50], [53]	10 eV	100 GeV	1	100	no
Brennan [60]	30 eV	700 keV	3	92	no
Chantler [30], [31]	10 eV	433 keV	1	92	K
Ebel [71]	1 keV	300 keV	1	92	all
Elam [67]	100 eV	1 MeV	1	98	no
EPDL97 [62]	10 eV	100 GeV	1	100	all
Henke [54]	10 eV	30 keV	1	92	no
McMaster [40], [42]	1 keV	700 keV	1	94	no
PHOTX [55]	1 keV	100 MeV	1	100	no
RTAB [66]	10 eV	300 keV	1	99	all
Scofield [29]	1 keV	1.5 MeV	1	100	all
StormIsrael [43]	1 keV	100 MeV	1	100	no
Veigele [45]	100 eV	1 MeV	1	94	no
XCOM-DB [58]	1 keV	500 keV	1	100	no
XCOM [58]	1 keV	100 GeV	1	100	no

they are identified in the validation analysis, the corresponding references and their coverage in energy, atomic number and atomic shells. A brief overview of their features is summarized in the following subsections, where the compilations appear in chronological order of publication; more details can be found the related references.

#### A. McMaster's Tables

The compilation by McMaster *et al.* [40], [41] reports coherent and incoherent photon scattering cross sections, photoelectric and total cross sections between 1 keV and 1 MeV. The data concern elements with atomic numbers 1 to 99, with the exception of atomic numbers 84, 85, 87, 88, 89, 91, and 93. The photoelectric cross sections were obtained by a combination of methods: by least-squares fitting of semi-empirical data, when they were available, by theoretical calculations from theory in regions where the data were insufficient or were considered unreliable, and by interpolation between semi-empirical and theoretical values, where no data of either kind were available. The semi-empirical photoelectric cross sections were derived by subtracting theoretical scattering cross sections from experimental total photon attenuation data. The theoretical cross sections are based on Schmickley and Pratt's [28] calculations.

The original compilation was updated in 2006 [42].

#### B. Storm and Israel's Tables

Storm and Israel's tables [43] encompass photon interaction cross sections for atoms with atomic numbers between 1 and 100 and photon energy from 1 keV to 100 MeV. The photoelectric cross sections are derived from several theoretical references: calculations by Brysk and Zerby [44], Rakavy and Ron [27], and Schmickley and Pratt [28].

#### C. Veigele's Tables

The photoelectric cross sections reported in Veigele's compilation [45] are of theoretical origin in the lower energy range (from 100 eV to energies varying between 1 and 10 keV,

depending on the element) and were calculated with semi-empirical methods at higher energies (up to 1 MeV).

The low energy cross sections derive from non-relativistic, self-consistent field calculations, which are based on independent particle approximation. At higher energies, theoretical scattering cross sections, calculated by relativistic self-consistent field methods, were subtracted from experimental total attenuation data; the resulting calculated photoelectric cross sections and available photoelectric cross section measurements from 1 keV to 1 MeV were then fitted by a least-squares procedure to produce recommended values.

Veigele's compilation concerns elements with atomic numbers from 1 to 94.

#### D. Scofield's Calculations

Scofield's 1973 [29] compilation reports photoeffect cross sections for all subshells, for all elements with atomic numbers from 1 to 101, over the photon energy range between 1 keV and 1.5 MeV. The calculations were later extended down to 100 eV [46].

These cross sections derive from non-relativistic calculations based on a solution of the Dirac equation for the orbital electrons moving in a static Hartree-Slater central potential. In this approximation the electron-electron interaction term is replaced with an average value, thus making the calculation less computationally intensive, but in principle also less accurate, than the full Hartree-Fock model, which requires the calculation of the self-consistent field for each term. In the same reference [29] Scofield reports renormalization factors for atomic numbers 2 to 54 to convert the cross sections calculated in the Dirac-Hartree-Slater approximation to values expected from a relativistic Hartree-Fock model.

Comparisons with experimental photon mass attenuation coefficients [46], [47] tend to favour Scofield's unrenormalized values over the renormalized ones. Nevertheless, it is worth noting that these evaluations of Scofield's photoeffect cross sections do not compare the two sets of calculations with direct photoelectric cross sections measurements: the reference data in these comparisons involve the subtraction of theoretically calculated photon scattering contributions from measured photon attenuation coefficients. It is also unclear whether these evaluations rest on a qualitative appraisal only or are based on objective statistical methods.

#### E. Biggs and Lighthill's Parameterisation

Biggs and Lighthill [48]–[50] expressed the total photoelectric cross sections as an empirical parameterization of data deriving from various semi-empirical and theoretical sources, which include, among others, Henke's 1982 compilation [51] and the 1978 version of EPDL [52], the latter in turn based on Storm and Israel's, McMaster's, Scofield's and Veigele's compilations.

For element  $i$  and energy range  $j$  the cross section is represented by the formula:

$$\sigma_{ij} = \frac{A_{ij1}}{E} + \frac{A_{ij2}}{E^2} + \frac{A_{ij3}}{E^3} + \frac{A_{ij4}}{E^4} \quad (1)$$



The tabulations in [50] list the most recent compilation of the  $A_{ij}$  coefficients of equation 1 assembled by the original authors of the parameterization.

Modified  $A_{ij}$  coefficients for some gases are documented in [53].

#### F. Henke's Compilation

The compilation of photoabsorption cross sections by Henke et al. [54] covers the energy range from 10 eV to 30 keV, for elements with atomic numbers up to 92.

Photoabsorption cross sections for energies below 10 keV are based on both theoretical calculations and experimental data, interpolating across the atomic number  $Z$  for elements where experimental data were scarce. Above 10 keV photoabsorption cross sections derive from the semi-empirical parameterizations by Biggs and Lighthill [50].

An earlier compilation by Henke [51], concerning a narrower energy range (from 30 eV to 10 keV), is based on similar criteria; cross sections from 1.5 to 10 keV are taken from [49].

#### G. PHOTX and XCOM

The PHOTX [55], [56] data library was developed as a basis for the photon cross section file of ENDF/B-VI [57]. It provides cross sections for coherent and incoherent scattering, photoelectric absorption, and pair production in the field of the nucleus and in the field of the atomic electrons. The data concern elements with atomic numbers between 1 and 100, and photon energies from 1 keV to 100 MeV.

XCOM [58] is a photon cross section database compiled by the NIST (National Institute of Standards and Technology of the United States of America). It concerns elements with atomic numbers up to 100 and photon energies from 1 keV to 100 GeV. Two sets of tabulations, identified as XCOM and XCOM-DB, are evaluated in this paper. XCOM corresponds to the standard energy grid available from the XCOM web site managed by NIST. XCOM-DB is encompassed in the DABAX database, which is part of the X-ray Optics Software Toolkit (XOP) [59]; it is mainly addressed to the photon science experimental community and is limited to photon energies up to 509.5 keV.

Regarding the photoelectric effect, both PHOTX and XCOM include Scofield's 1973 unrenormalized cross sections [29] up to 1.5 MeV. At higher energies a semi-empirical formula from [32] connects Scofield's values at 1.5 MeV to the asymptotic high energy limit calculated by Pratt [25].

#### H. Brennan and Cowan's Calculations

Brennan and Cowan developed a collection of software programs [60] for the calculation of photoabsorption cross sections, atomic scattering factors and other quantities relevant to photon science. The photoelectric cross sections are based on Cromer and Liberman's [61] theory.

Tabulations of photoelectric cross sections derived from Brennan and Cowan's calculations are included in the XOP software system [59].

#### I. EPDL

The EPDL (Evaluated Photon Data Library) includes photon interaction data concerning photoionization, photoexcitation, coherent and incoherent scattering, and pair and triplet production. The latest version at the time of writing this paper was released in 1997, and is commonly identified as EPDL97 [62]. It is part of the ENDF/B-VII.1 [63] evaluated nuclear data file.

EPDL97 includes total and partial cross sections for elements with atomic numbers between 1 and 100, and for photon energies from 1 eV to 100 GeV. Partial cross sections are tabulated for all subshells.

From the edge to 1 MeV, subshell ionization cross sections are based on Scofield's data as in [46]; total photoionization cross sections are summed over all subshells. From 1 MeV to 100 GeV the total cross sections are based on Hubbell's data reported in [64]. Scofield's subshell cross sections have been extended up to 100 GeV by ensuring that the sum of the subshell cross sections is equal to Hubbell's total, and maintaining the same ratio between subshell cross sections over the entire energy range from 1 MeV to 100 GeV. At 1 MeV the total photoionization cross section is identical from both sources, therefore the two sets of calculations could be combined in a consistent manner.

EPDL97 documentation reports rough estimates and qualitative comments about the accuracy of the tabulated data, but it does not document how these estimates were produced. To the best of our knowledge systematic, quantitative validation of EPDL97 photoionization data is not documented in the literature.

EPDL97 is extensively used in Monte Carlo simulation; details are given in Section V.

#### J. Chantler's Calculations

Chantler calculated photoelectric cross sections in a self-consistent relativistic Dirac-Hartree-Fock framework [30], [31]. The exchange potential of Chantler's approach follows that of Cromer and Liberman [61], [65] and Brennan and Cowan [60], and is different from the approach used by Scofield [29].

The calculations are carried out in isolated atom and independent particle approximations: each atom is treated as a standalone system, not influenced by other atoms or particles, and each electron is considered to move in an effective potential of the nucleus with the average repulsive force of the electrons. This effective screening neglects some correlation; it also neglects the fact that the potential for one electron is not identical to that of a different electron.

Chantler's published tabulations report total and K-shell cross sections for elements with atomic number from 1 to 92; they cover an energy range comprised between 1–10 eV and 0.4–1 MeV (the lower and upper bounds vary with the atomic number).

#### K. RTAB

The RTAB [66] database encompasses three sets of photoionization cross sections along with an extensive set of tabulations

concerning photon elastic scattering: a set of original calculations (identified in the following as RTAB cross sections), an extension of these cross sections based on EPDL97 (identified as RTABX) and Scofield's 1973 cross sections. They are listed for atomic numbers from 1 to 100.

All the data in the RTAB database (apart from those reported from other sources) have been consistently computed in the same Dirac-Slater potential. Scofield's 1973 [29] cross sections were also computed in a Dirac-Slater potential, although differently from those included in the RTAB database. According to [66], subshell cross sections are calculated in the Dirac-Slater potential to obtain total atom cross sections from threshold up to several hundred keV. For higher energies, these values are smoothly joined to the total-atom values in the EPDL97, thus extending the cross sections in the RTABX collection up to 100 GeV. RTAB photoelectric cross sections are tabulated for energies up to 300 keV.

The RTAB photoelectric database has not been exploited yet in general purpose Monte Carlo systems.

#### L. Elam's Database

Elam's database [67] provides photon cross section data for elements with atomic numbers between 1 and 98, and for photon energy from 100 eV to 1 MeV.

Above 1 keV the photoabsorption cross sections derive from the XCOM ones of [68], which were parameterized with a cubic spline algorithm. Below 1 keV they are based on the 1981 version of EPDL [69]. Appropriate algorithms were applied to the data to ensure a smooth connection between the two sets of cross sections. Edge discontinuities were adjusted to be consistent with Williams' atomic binding energies [70].

#### M. Ebel's Parameterizations

Ebel *et al.* [71] developed parameterizations of total photoelectric absorption coefficients and of subshell absorption coefficients in the energy range from 1 keV to 300 keV, for elements with atomic number up to 92. They are based on fitting fifth order polynomials in the logarithm of the photon energy to Scofield's 1973 [29] cross section data. The coefficients for the parameterization are tabulated for each energy interval identified by the absorption edges of a given element. Photoelectric cross sections can be calculated using the tabulations.

This compilation was developed especially for application in fundamental parameter programs for quantitative X-ray analysis.

## V. PHOTOIONIZATION IN MONTE CARLO CODES

General purpose Monte Carlo codes for particle transport include algorithms for the simulation of the photoelectric effect.

The original version of EGS4 [72] calculated photoelectric total cross sections based on Storm and Israel's tables [43] and generated the photoelectron with the same direction as the incident photon. Later evolutions introduced the use of PHOTX cross sections [73] and the generation of the photoelectron angular distribution [74] based on Sauter's theory [34]. These

features are currently implemented in EGS5 [14]. EGSnrc [75] provides the option of calculating total photoelectric cross sections based on Storm and Israel's tables as originally in EGS4 or on a fit to XCOM [58] cross sections, while it uses subshell cross sections based on EPDL [62]. It samples the photoelectron angular distribution according to the method described in [74] based on Sauter's theory.

ETRAN [76] uses Scofield's 1973 [29] cross sections for energies from 1 keV to 1.5 MeV and extends them to higher energies by exploiting Hubbell's method [32] to connect the values at 1.5 MeV to the asymptotic high energy limit calculated by Pratt [25]. It samples the direction of the photoelectron from Fischer's [33] distribution for electron energies below 50 keV and from Sauter's [34] distribution for higher energies.

FLUKA [77], [78] calculates photoelectric cross sections based on EPDL97 and samples the photoelectron direction according to Sauter's theory [34].

ITS [79] calculates photoelectric cross sections based on Scofield's 1973 non-renormalized values. The angle of the photoelectron with respect to the parent photon is described by Fischer's distribution [33] at lower energies and by Sauter's [34] formula at higher energies.

MCNP5 [80] and MCNPX [81] provided different options of data libraries for the calculation of photoelectric cross sections: two version of EPDL (EPDL97 [62] and EPDL89 [82]), and ENDF/B-IV [84] data complemented by Storm and Israel's tables [43] for atomic numbers greater than 83. MCNP6 [83] has extended the minimum energy cut-off for photon transport down to 1 eV; the necessary photo-atomic cross sections derive from ENDF/B-VI version 8, which in turn is based on EPDL97 regarding photon interactions.

In the first version of Penelope including photon transport [85] photoelectric cross sections were interpolated from XCOM; in more recent versions [86], [87] they are interpolated from EPDL97 tabulations. The photoelectron angular distribution is sampled from Sauter's differential cross section for the K shell [34].

GEANT 3 [88] calculated total photoionization cross sections based on Biggs and Lighthill's [50] parameterizations; the probability of ionization of the K shell and L subshells was estimated by parameterizations of the jump ratios deriving from Veigele's [45] tables. The angular distribution of the photoelectron was sampled for the K shell and for the  $L_1$ ,  $L_2$  and  $L_3$  subshells based on Sauter's [34], [35] and Gavrila's [37], [39] calculations.

The Geant4 toolkit encompasses various implementations of the photoelectric effect. The overview summarized here concerns the latest version at the time of writing this paper: Geant4 10.1, complemented by two correction patches.

The model implemented in Geant4 *standard* electromagnetic package calculates cross sections based on the analytical formula of Biggs and Lighthill, but it uses modified coefficients deriving from a fit to experimental data. The related reference cited in Geant4 Physics Reference Manual [89] as the source of these modifications does not appear to be consistent, presumably due to a mismatch between the Russian and English versions of the periodical where it was published. The modifications appear to derive from [53], which reports fits

to experimental data concerning noble gases, hydrogen, carbon, fluorine, oxygen and silicon; they concern cross sections at low energies. The atomic cross section calculation according to Biggs-Lighthill parameterization is implemented in the *G4SandiaTable* and *G4StaticSandiaData* classes of Geant4 *materials* package. The same calculation of photoelectric cross sections is also used by the PAI (PhotoAbsorption-Ionisation) model [90]. The energy of the emitted photoelectron is determined as the difference between the energy of the interacting photon and the binding energy of the ionized shell defined in the *G4AtomicShells* class [22], and the photoelectron angle is calculated according to the Sauter-Gavrila distribution for K shell [34], [37].

Geant4 *low energy* electromagnetic package [91], [92] encompasses two implementations of the photoelectric effect, one identified as “Livermore” [93] and one reengineered from the 2008 version of the Penelope code [86]; both models calculate total and partial cross sections based on EPDL97. The so-called “Livermore” model provides three options of computing the angular distribution of the emitted photoelectron: in the same direction as the incident photon, based on Gavrila’s distribution of the polar angle [37] for the K shell and the  $L_1$  subshell, and based on a double differential cross section derived from Gavrila’s [37], [39] calculations, which can also handle polarized photons.

In addition, the Geant4 toolkit encompasses two models for the simulation of the photoelectric effect concerning polarized photons: one for circularly polarised photons in the *polarisation* package and one in the *low energy* electromagnetic package, identified as “Livermore polarized”. Their evaluation is outside the scope of this paper.

## VI. THE VALIDATION PROCESS

The validation process adopts the same methodology used in [94] for the validation of photon elastic scattering. The main concepts and most relevant features of the validation method are summarized below to facilitate understanding of the results reported in Sections VII and VIII.

### A. Simulation Models

The validation process concerns the methods for the calculation of total and subshell cross sections summarized in Table I. Some of these simulation models represent novel approaches with respect to those so far available in Geant4 and in other general purpose Monte Carlo codes.

All the physics models subject to evaluation have been implemented in a consistent software design, compatible with the Geant4 toolkit, which minimizes external dependencies to ensure unbiased appraisal of their intrinsic behaviour. The software adopts a policy-based class design [95], which supports the provision of a wide variety of physics models without imposing the burden of inheritance from a predefined interface.

A single policy class calculates cross sections that exploit tabulations; alternative tabulations, corresponding to different physics models, are managed through the file system. Dedicated policy classes implement cross section calculations based on

analytical formulae. The same scheme is adopted for total and partial cross section calculation; the latter is involved in the algorithm that creates a vacancy, which drives the subsequent atomic relaxation process. Alternative modeling of photoelectron angular distributions is handled through a Strategy pattern [100].

A photoionization process, derived from the *G4VDiscreteProcess* class of Geant4 kernel, which in turn is derived from *G4VProcess*, acts as a host class for the policy classes; they can be interchanged [96]–[98] to determine its behaviour. The simulation of photoionization according to this software design is consistent with Geant4 kernel, since Geant4 tracking handles all processes polymorphically through the *G4VProcess* base class.

Since policy classes are characterized by a single responsibility and have minimal dependencies on other parts of the software, the adopted programming paradigm enables independent modeling and test of all physics options. Their validation can be performed through simple unit tests. This strategy ensures greater modeling flexibility and testing agility than the one adopted in the Geant4 electromagnetic package, where total cross section, vacancy creation and final state generation are bundled into one object: that software design choice requires full-scale simulation applications to test basic physics entities, due to extensive dependencies imposed by the *G4VEmModel* base class, from which all photoelectric models derive.

Existing physics models in Geant4 have been refactored [101] consistently with this software design; other models not yet available in Geant4 have been implemented for the first time. The correctness of the implementation has been verified prior to the validation process to ensure that the software reproduces the physical features of each model consistently.

### B. Experimental Data

Experimental data [102]–[181] for the validation of the simulation models were collected from a survey of the literature. The sample of experimental total cross sections consists of approximately 3000 measurements, which concern 61 target atoms and span energies approximately from 5 eV to 1.2 MeV. It includes measurements at energies below 100 eV, mostly concerning gaseous targets: these data are relevant to evaluate the accuracy of calculations performed in independent particle approximation at very low energies, e.g. the EPDL97 data library, which extend down to 1 eV.

The sample of subshell cross sections encompasses approximately 600 measurements, which concern 52 target atoms and span energies approximately from 1 eV to 2.75 MeV.

An overview of the experimental data sample is summarized in Tables II–III.

The photoionization cross sections reported in the literature as experimentally measured often derive from measurements of total photon attenuation, from which theoretically calculated contributions from photon scattering were subtracted. These semi-empirical values are not appropriate to an epistemologically correct validation process, which requires the comparison of simulation models with truly experimental data. An evaluation of the possible systematic effects induced by using

TABLE II  
SUMMARY OF THE EXPERIMENTAL TOTAL CROSS SECTION DATA USED IN THE VALIDATION ANALYSIS

Atomic Number	Element	Energy range (keV)	Sample size	References
1	H	0.0136 - 0.020	27	[102]–[104]
2	He	0.025 - 0.277	320	[105]–[112]
3	Li	0.046 - 0.400	93	[113]–[116]
7	N	0.015 - 0.4	73	[117], [118]
8	O	0.013 - 0.28	215	[103], [119]–[121]
10	Ne	0.022 - 2.952	448	[108], [110]–[112], [117], [122]
11	Na	0.046 - 0.246	17	[123]
13	Al	145.4	1	[124]
17	Cl	0.016 - 0.078	25	[125]
18	Ar	0.016 - 6	487	[105], [110]–[112], [117], [126]–[130]
19	K	0.004 - 0.005	12	[131]
22	Ti	59.54	1	[132]
23	V	59.54	1	[132]
24	Cr	59.54	1	[132]
25	Mn	59.54	1	[132]
26	Fe	0.008 - 59.54	25	[132]–[134]
27	Co	59.54	1	[132]
28	Ni	1.487 - 59.54	17	[132], [133]
29	Cu	59.54 - 661.6	9	[124], [132], [135]–[138]
30	Zn	59.54	1	[132]
33	As	59.54	1	[132]
34	Se	59.54	1	[132]
36	Kr	0.015 - 1.626	357	[105], [110], [112], [139], [140]
37	Rb	0.004 - 0.010	4	[141]
38	Sr	59.54	1	[132]
39	Y	279.2 - 661.6	2	[142]
40	Zr	59.54 - 661.6	7	[124], [135]–[137], [143]
41	Nb	59.54	2	[143]
42	Mo	59.54 - 661.6	7	[132], [138], [143]
43	Tc	59.54	1	[143]
44	Ru	59.54	1	[143]
45	Rh	59.54	1	[143]
46	Pd	59.54	1	[143]
47	Ag	1.487 - 661.6	13	[124], [?], [135]–[137], [143]–[146]
48	Cd	59.54	2	[132], [143]
49	In	59.54	2	[143]
50	Sn	1 - 661.6	31	[124], [133], [135]–[137], [143], [144], [146]
51	Sb	59.54	1	[143]
52	Te	59.54	2	[143]
54	Xe	0.013 - 6	657	[110], [122], [128], [130], [147]
55	Cs	0.004 - 59.54	14	[132], [141], [148]
57	La	59.54	1	[149]
58	Ce	59.54 - 661.6	6	[142], [149]–[152]
59	Pr	59.54 - 661.6	3	[150]–[152]
60	Nd	59.54 - 661.6	2	[150], [152]
61	Pm	59.54	1	[152]
62	Sm	59.54 - 145.4	4	[149], [151], [152]
63	Eu	59.54	2	[149], [152]
64	Gd	59.54 - 661.6	5	[150]–[153]
65	Tb	59.54	3	[149], [152], [153]
66	Dy	59.54 - 661.6	8	[142], [149]–[153]
67	Ho	59.54	1	[152]
68	Er	59.54	3	[149], [152], [153]
70	Yb	279.2 - 661.6	2	[142]
72	Hf	65.839–68.547	25	[154]
73	Ta	1.487 - 661.6	31	[124], [133], [135]–[138], [144]
78	Pt	1 - 40	25	[133]
79	Au	0.93 - 661.6	33	[133], [135]–[137], [145]
80	Hg	1173	1	[155]
82	Pb	1.487 - 1173	36	[124], [133], [135]–[137], [144]–[146], [155], [156]
92	U	1.487 - 3	4	[133]

semi-empirical data [182]–[245] as a reference for testing cross section calculations is reported in Section VII-A; it concerns a sample of approximately 1500 total cross sections, spanning energies between approximately 50 eV and 6 MeV.

Some experimental measurements have been published only in graphical form; numerical values were extracted from the

plots by means of the PlotDigitizer [246] digitizing software. The error introduced by the digitization process was estimated by digitizing a few experimental data samples, which are reported in the related publications both in graphical and numerical format. The reliability of the digitized values is hindered by the difficulty of appraising the experimental points



TABLE III  
SUMMARY OF THE EXPERIMENTAL SUBSHELL CROSS SECTION DATA USED IN THE VALIDATION ANALYSIS

Z	Element	Shell	Energy range (keV)	Sample size	References
3	Li	K	0.069-1.487	10	[115]
10	Ne	L <sub>1</sub>	0.006-0.127	14	[158]
18	Ar	M <sub>1</sub>	0.001-0.1	147	[159]–[162]
20	Ca	M <sub>1</sub> , N <sub>1</sub>	0.038-0.121	24	[157]
25	Mn	M <sub>1</sub>	0.113-0.27	7	[163]
26	Fe	K	59.54	1	[164]
28	Ni	K	59.54	1	[164]
29	Cu	K	59.54-441.8	3	[164], [165]
30	Zn	K	59.54	1	[164]
33	As	K	59.54	1	[164]
34	Se	K	59.54	1	[164]
35	Br	K	59.54	1	[164]
36	Kr	N <sub>1</sub>	0.028-0.041	5	[162]
40	Zr	K, L <sub>1</sub> -L <sub>3</sub>	59.54-441.8	15	[124], [164]–[166]
41	Nb	K, L, L <sub>1</sub> -L <sub>3</sub>	59.54	6	[164], [166]
42	Mo	K, L, L <sub>1</sub> -L <sub>3</sub>	5.96-74.4	6	[164], [166]
43	Tc	K, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
44	Ru	K, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
45	Rh	K, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
46	Pd	K, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
47	Ag	K, L, L <sub>1</sub> -L <sub>3</sub>	59.54-441.8	8	[124], [164]–[166]
48	Cd	K, L, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
49	In	K, L, L <sub>1</sub> -L <sub>3</sub>	59.54	6	[164], [166]
50	Sn	K, L, L <sub>1</sub> -L <sub>3</sub>	59.54-1330	19	[124], [164]–[166], [167], [168], [169]
51	Sb	K, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
52	Te	K, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
53	I	K	59.54	2	[164]
54	Xe	M <sub>4</sub> , M <sub>5</sub> , O <sub>1</sub>	0.0236-1	66	[162], [170], [171]
56	Ba	O <sub>1</sub>	0.069-0.119	28	[172]
58	Ce	K, L, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
59	Pr	K, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
60	Nd	K, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
61	Pm	K, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
62	Sm	K, L, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
63	Eu	K, L, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
64	Gd	K, L <sub>1</sub> -L <sub>3</sub>	50.3-59.54	5	[166], [173]
65	Tb	K, L, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
66	Dy	K, L, L <sub>1</sub> -L <sub>3</sub>	53.8-59.54	5	[166], [174]
67	Ho	K, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
68	Er	K, L, L <sub>1</sub> -L <sub>3</sub>	59.54	4	[166]
70	Yb	K	61.34	1	[174]
72	Hf	K	65.29-65.44	2	[173], [175]
73	Ta	K	67.36-1330	15	[124], [165], [167]–[169], [173], [175]
74	W	K	59.57-69.64	2	[174], [176]
78	Pt	K	1330	1	[155]
79	Au	K, N <sub>6</sub> -N <sub>7</sub> , O <sub>1</sub> -O <sub>3</sub> , P <sub>1</sub>	5-1330	60	[26], [165], [167], [169], [173], [175], [177]
80	Hg	K, L	59.54-1250	2	[155], [176]
81	Tl	L	59.54-74.4	3	[176], [178]
82	Pb	K, L, L <sub>3</sub>	13.6-2750	28	[26], [124], [155], [165], [167]–[169], [173], [175], [176], [178]–[180]
83	Bi	K, L	59.54-1330	4	[155], [176], [178]
90	Th	K, L <sub>3</sub>	16.9-1330	9	[167], [169], [179]
92	U	K, L, L <sub>3</sub>	17.8-1330	10	[155], [176], [178], [179], [181]

and their error bars in figures that may span several orders of magnitude in logarithmic scale, or that appear of questionable graphical quality in the original publication. Caution was exercised in using these digitized data in the validation analysis; they were discarded, if incompatible with other measurements reported in the literature in numerical form.

Large discrepancies are evident in some of the experimental data; systematic effects are probably present in some cases, where sequences of positive or negative differences between data samples originating from different experimental groups are qualitatively visible, and confirmed by the Wald-Wolfowitz test [247] to be incompatible with randomness. Experimental data exhibiting large discrepancies with respect to

other measurements in similar configurations, which hint at the presence of systematic effects, are excluded from the validation tests.

The validation process is hindered by physical effects related to the structure of the target material. Accuracy of edge position is limited by chemical shifts and the detailed structure of the experimental material observed. Usually an accuracy of absolute energies below 1–3 eV is unattainable for this reason. At low energies (less than 200–500 eV) the occurrence of collective valence effects and dipole resonances can lead to much larger deviations (e.g. up to 50 eV or 10%). General purpose Monte Carlo codes do not take into account such material structure effects; the cross sections they use for the simulation of



the photoionization process, briefly outlined in Section IV, are not intended to model these features. This limitation should be taken into account in the evaluation of the results of the validation process.

Correct estimate of experimental errors is a concern in the validation of physics models, since unrealistic estimation of the experimental errors may lead to incorrect conclusions regarding the rejection of the null hypothesis in tests whose statistic takes experimental uncertainties explicitly into account. Although technological developments have contributed to improved precision of measurement, some estimates of experimental uncertainties reported in the literature may be excessively optimistic, especially when they appear inconsistent with other measurements exploiting similar experimental techniques. Experimental measurements claiming much smaller uncertainties than similar ones have been critically evaluated in the analysis process.

### C. Data Analysis

The evaluation of the simulation models performed in this study has two objectives: to validate them quantitatively, and to compare their relative capabilities.

The scope of the software validation process is defined according to the guidelines of the IEEE Standard devoted to software verification and validation [248], which conforms to the life cycle process standard defined in ISO/IEC Standard 12207 [249]. For the problem domain considered in this paper, the validation process provides evidence that the software models photoionization consistently with experiment.

A quantitative analysis, based on statistical methods, is practically possible only for the validation of cross sections, for which a large sample of experimental data is available. The scarcity of angular distribution data in the literature hinders the validation of simulation models through similar statistical analysis methods: only qualitative general considerations can be made.

The statistical analysis of photoionization cross sections is articulated over two stages: the first determines the compatibility between the cross sections calculated by each simulation model and experimental data, while the second determines whether the various models exhibit any significant difference in compatibility with experiment. The Statistical Toolkit [250], [251] and R [252] are used in the statistical analysis. The level of significance of the tests is 0.01, unless stated otherwise.

The first stage of the analysis encompasses a number of test cases, each one corresponding to a configuration (characterized by photon energy, target element, experimental source and, if appropriate, subshell) for which experimental data are available. The inclusion of the experimental source in the definition of a test case facilitates the identification of possible systematic effects related to the experimental environment of the measurements. For each test case, cross sections calculated by the software are compared with measured ones by means of goodness-of-fit tests. The null hypothesis is defined as the equivalence of the simulated and experimental data distributions subject to comparison, as being drawn from the same parent distribution. The goodness-of-fit analysis is primarily

based on the  $\chi^2$  test [253]. Among goodness-of-fit tests, this test has the property of taking experimental uncertainties explicitly into account; consequently, the test statistic is sensitive to their correct appraisal.

The “efficiency” of a physics model is defined as the fraction of test cases in which the  $\chi^2$  test does not reject the null hypothesis. This variable quantifies the capability of that simulation model to produce results statistically consistent with experiment over the whole set of test cases, which in physical terms means over the whole range of photon energies and target elements involved in the validation process. Two methods were applied to calculate the uncertainties on the efficiencies: the conventional method involving the binomial distribution, described in many introductory statistics textbooks (e.g. [254]), and a method based on Bayes’ theorem [255]. The two methods deliver identical results within the number of significant digits reported in the following tables; the method based on Bayes’ theorem delivers meaningful results, which are reported in the following sections, also in limiting cases, i.e. for efficiencies very close to 0 or to 1, where the conventional method based on the binomial distribution produces unreasonable values.

The second stage of the statistical analysis quantifies the differences of the simulation models in compatibility with experiment. It consists of a categorical analysis based on contingency tables, which derive from the results of the  $\chi^2$  test: the outcome of this test is classified as “fail” or “pass”, according to whether the null hypothesis is rejected or not, respectively. The simulation model exhibiting the largest efficiency is considered as a reference in the categorical analysis; the other models are compared to it, to determine whether their difference in compatibility with measurements is statistically significant.

The null hypothesis in the analysis of a contingency table assumes equivalent compatibility with experiment of the cross section models it compares.

A variety of tests is applied to determine the statistical significance of the difference between the data categories (i.e. cross section models) subject to evaluation: Pearson’s  $\chi^2$  test [256] (when the number of entries in each cell of the table is greater than 5), Fisher’s exact test [257], Boschloo’s [258] test, the test based on Suissa and Schuster’s Z-pooled statistic [259], Santner and Snell’s test [260] and Barnard’s test [261]. As some contingency tables contain cells with a large number of entries ( $> 100$ ), Barnard’s test was calculated according to the approximate CSM statistic [262] to reduce the computational burden. The use of a variety of tests mitigates the risk of introducing systematic effects in the validation results due to peculiarities in the mathematical formulation of the test statistic.

Fisher’s test is widely used in the analysis of contingency tables. It is based on a model in which both the row and column sums are fixed in advance, which seldom occurs in experimental practice; it remains applicable to cases in which the row or column totals, or both, are not fixed, but in these cases it tends to be conservative, yielding a larger p-value than the true significance of the test [264].

Unconditional tests, such as Barnard’s test [261], Boschloo’s test [258] and Suissa and Shuster’s [259] calculation of a Z-pooled statistic, are deemed more powerful than Fisher’s

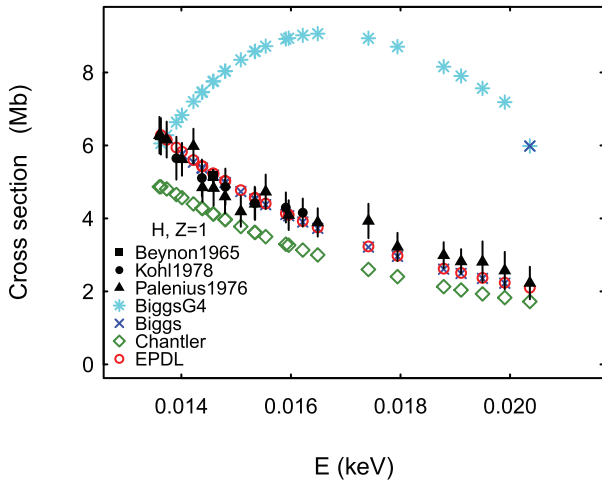


Fig. 1. Total photoionization cross section for  $Z = 1$  as a function of photon energy.

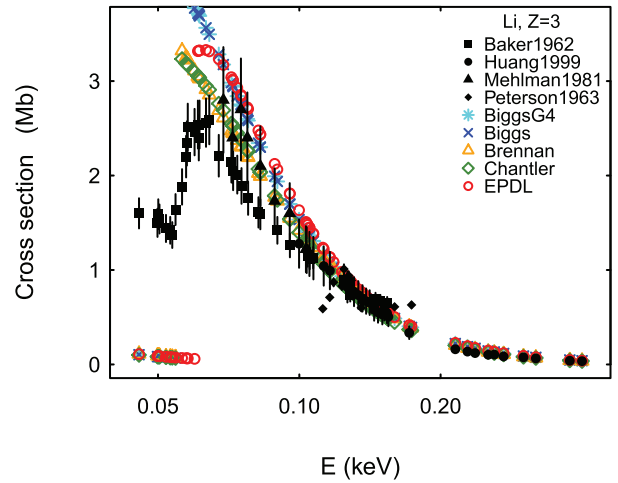


Fig. 3. Total photoionization cross section for  $Z = 3$  as a function of photon energy.

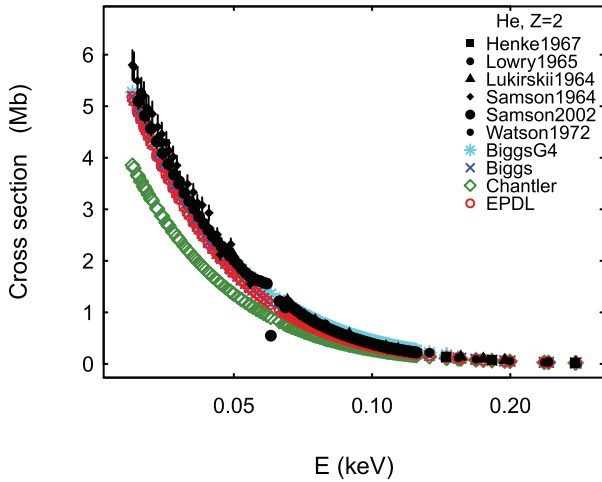


Fig. 2. Total photoionization cross section for  $Z = 2$  as a function of photon energy.

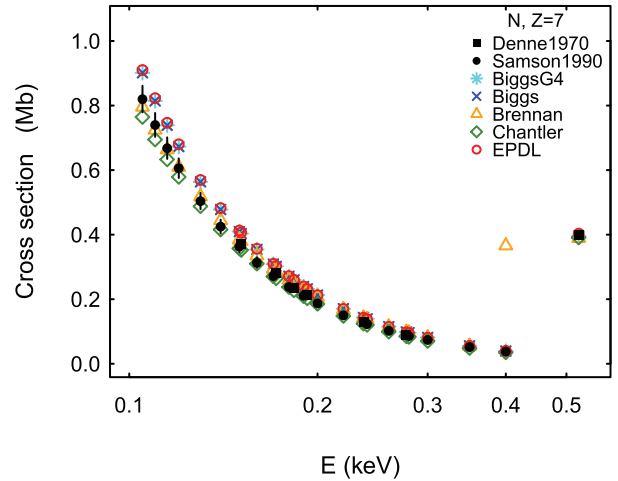


Fig. 4. Total photoionization cross section for  $Z = 7$  as a function of photon energy.

exact test in some configurations of  $2 \times 2$  contingency tables [265], [266], but they are computationally more intensive.

### VII. RESULTS OF TOTAL CROSS SECTION VALIDATION

Figs. 1 to 24 illustrate calculated and experimental total cross sections.

The validation analysis encompasses various areas of investigation: the evaluation of possible systematic effects related to the characteristics of reference data, the evaluation of cross section models covering a wide energy range, the evaluation of specialized models with limited energy coverage and the appraisal of the capability of the examined cross section calculations to describe the photoelectric effect at the low energy end.

#### A. Evaluation of Systematic Effects Related to Reference Data

An analysis was performed prior to the proper validation process to establish whether semi-empirical cross sections could be

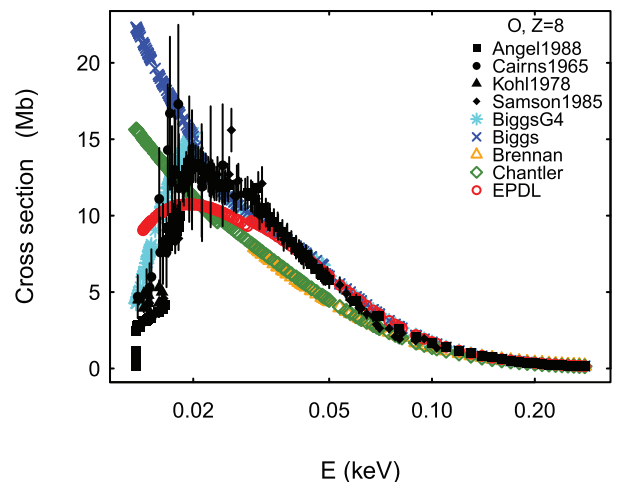


Fig. 5. Total photoionization cross section for  $Z = 8$  as a function of photon energy.

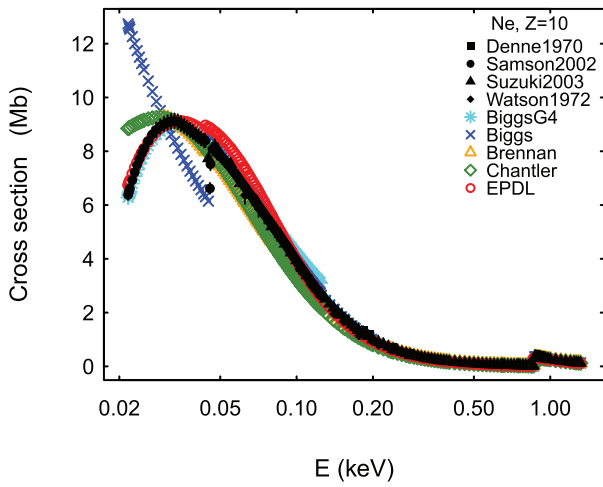


Fig. 6. Total photoionization cross section for  $Z = 10$  as a function of photon energy.

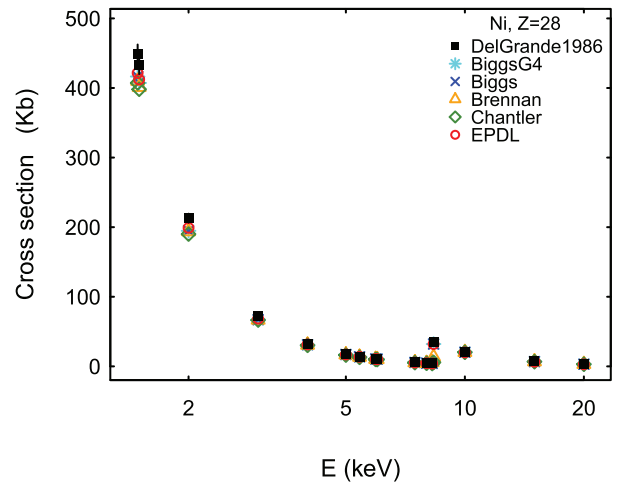


Fig. 9. Total photoionization cross section for  $Z = 28$  as a function of photon energy.

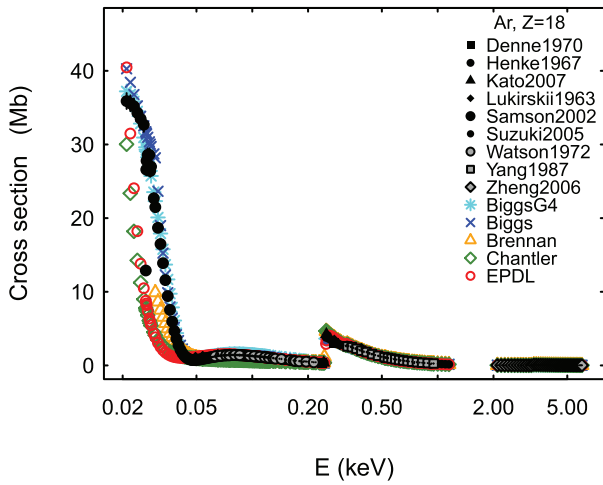


Fig. 7. Total photoionization cross section for  $Z = 18$  as a function of photon energy.

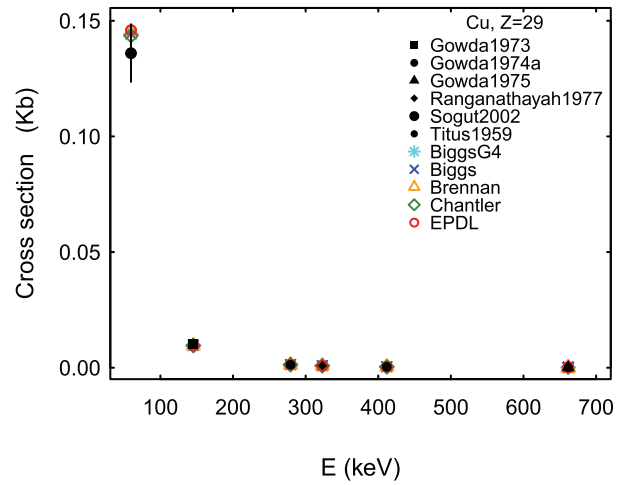


Fig. 10. Total photoionization cross section for  $Z = 29$  as a function of photon energy.

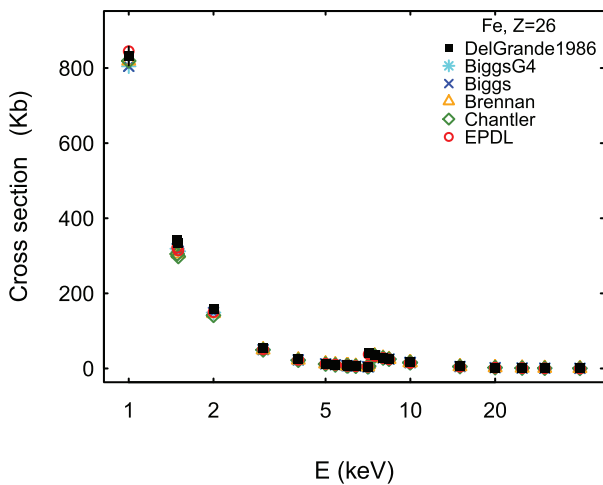


Fig. 8. Total photoionization cross section for  $Z = 26$  as a function of photon energy.

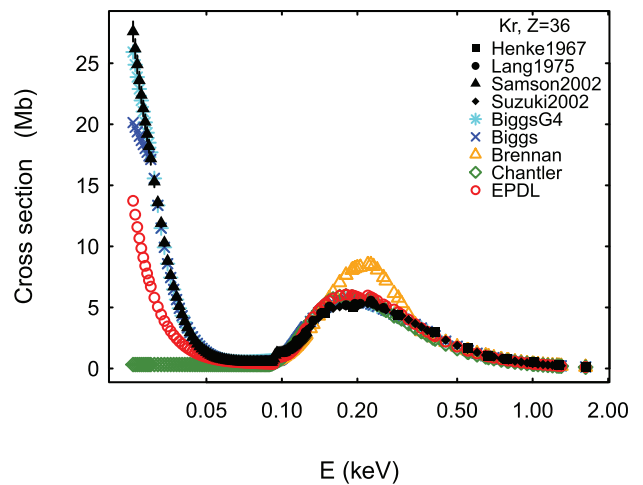


Fig. 11. Total photoionization cross section for  $Z = 36$  as a function of photon energy.

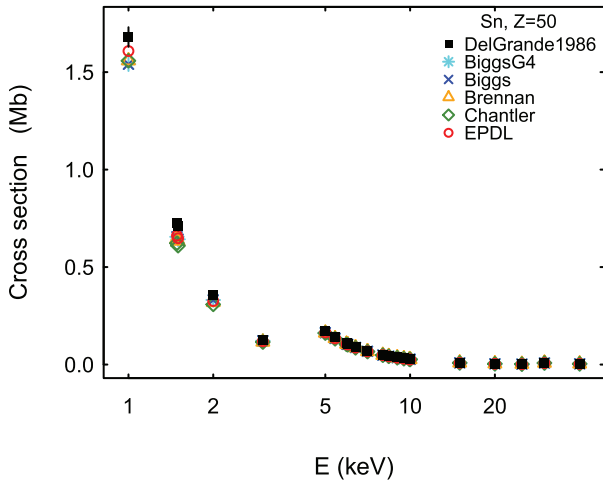


Fig. 12. Total photoionization cross section for  $Z = 50$  as a function of photon energy.

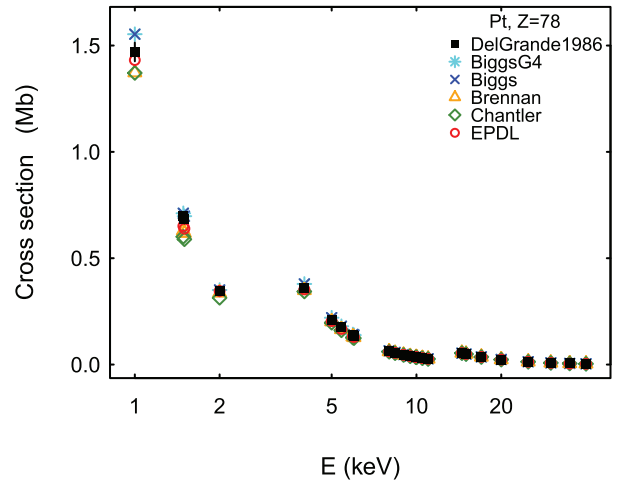


Fig. 15. Total photoionization cross section for  $Z = 78$  as a function of photon energy.

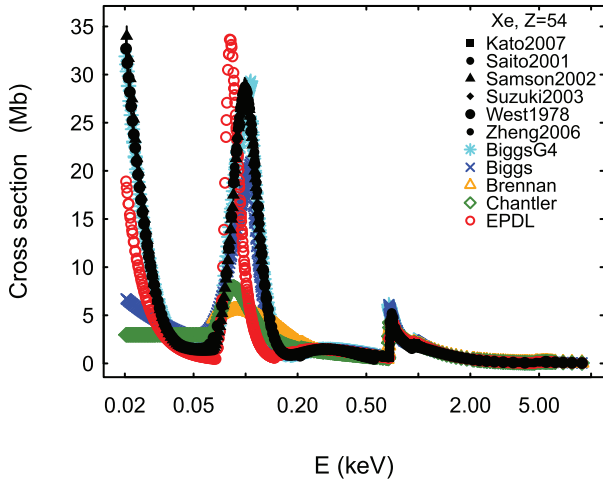


Fig. 13. Total photoionization cross section for  $Z = 54$  as a function of photon energy.

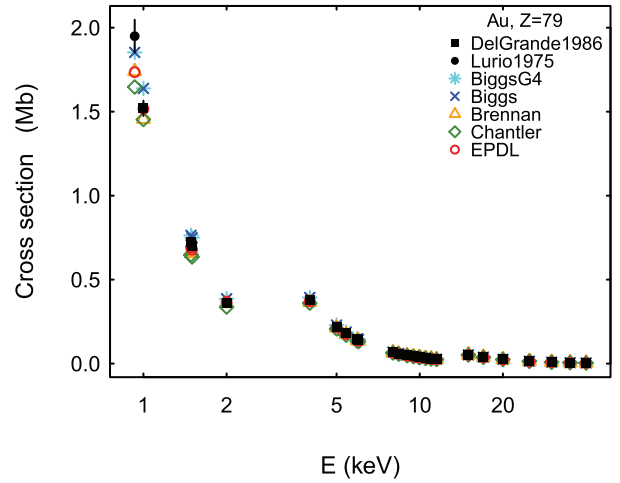


Fig. 16. Total photoionization cross section for  $Z = 79$  as a function of photon energy.

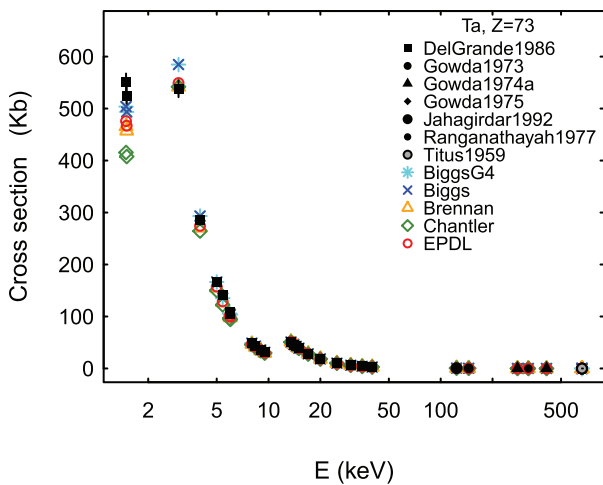


Fig. 14. Total photoionization cross section for  $Z = 73$  as a function of photon energy.

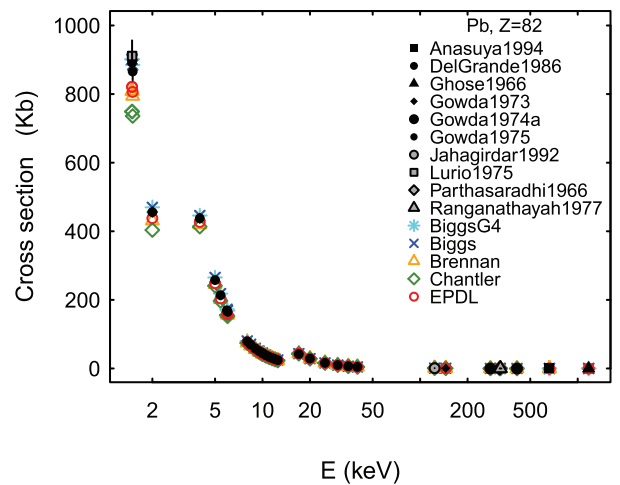


Fig. 17. Total photoionization cross section for  $Z = 82$  as a function of photon energy.



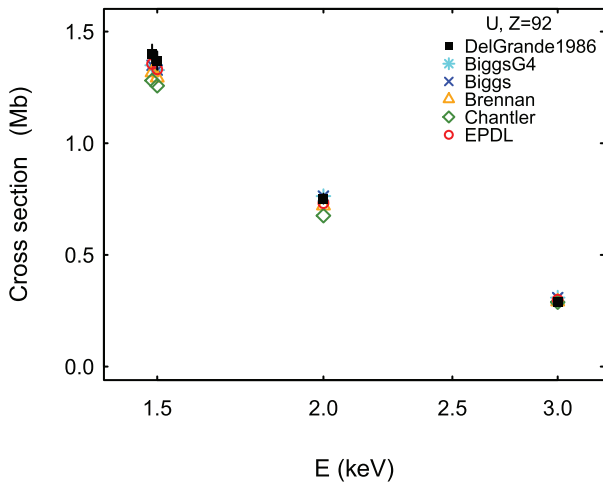


Fig. 18. Total photoionization cross section for  $Z = 92$  as a function of photon energy.

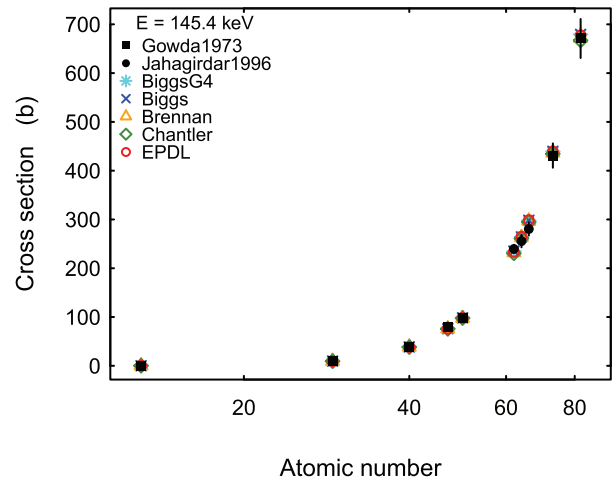


Fig. 21. Total photoionization cross section at 145.4 keV as a function of the atomic number  $Z$ .

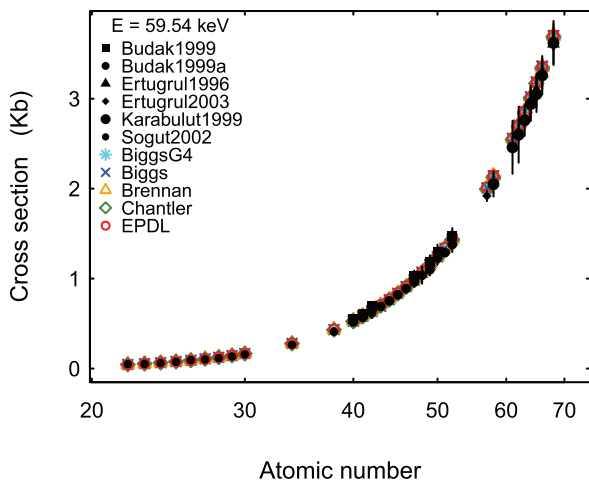


Fig. 19. Total photoionization cross section at 59.54 keV as a function of the atomic number  $Z$ .

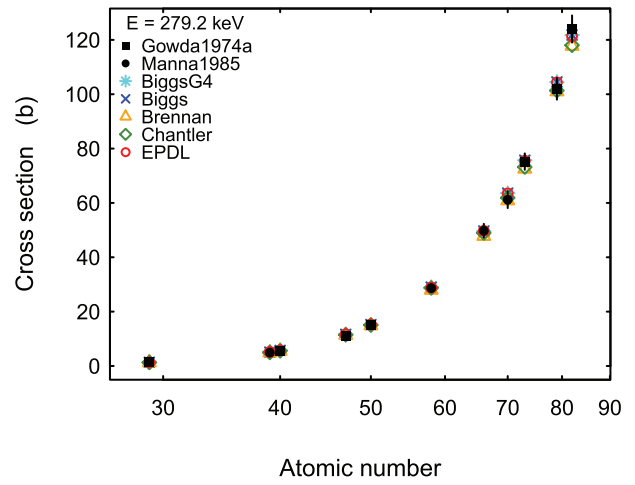


Fig. 22. Total photoionization cross section at 279.2 keV as a function of the atomic number  $Z$ .

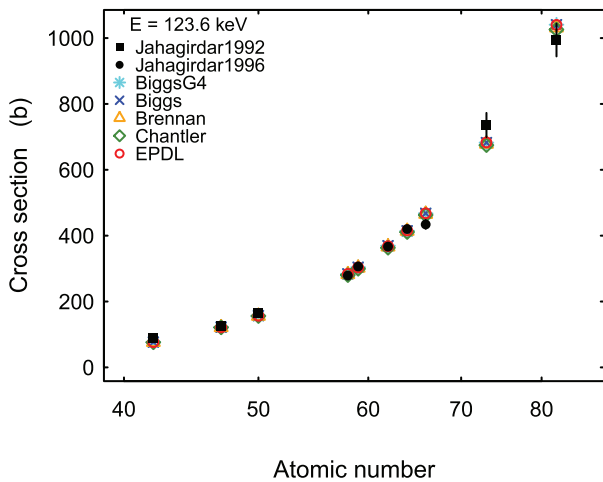


Fig. 20. Total photoionization cross section at 123.6 keV as a function of the atomic number  $Z$ .

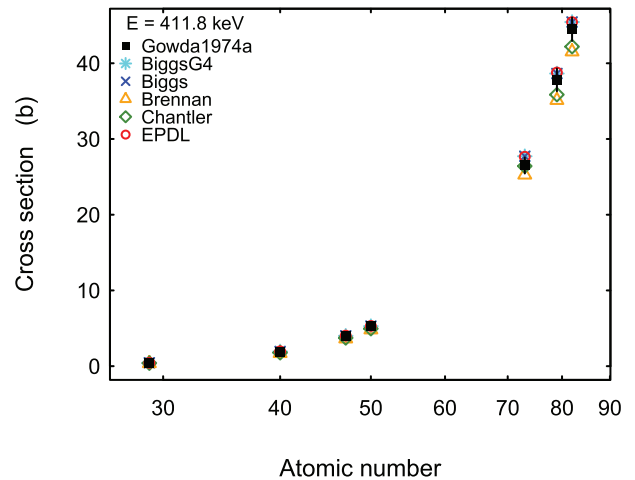


Fig. 23. Total photoionization cross section at 411.8 keV as a function of the atomic number  $Z$ .

used as a reference for comparison without introducing systematic effects. This evaluation concerns photon energies greater than 1 keV, since most of the semi-empirical data collected from

the literature are above this energy; it involves cross section compilations that are applicable over the whole energy covered by experimental and semi-empirical data.

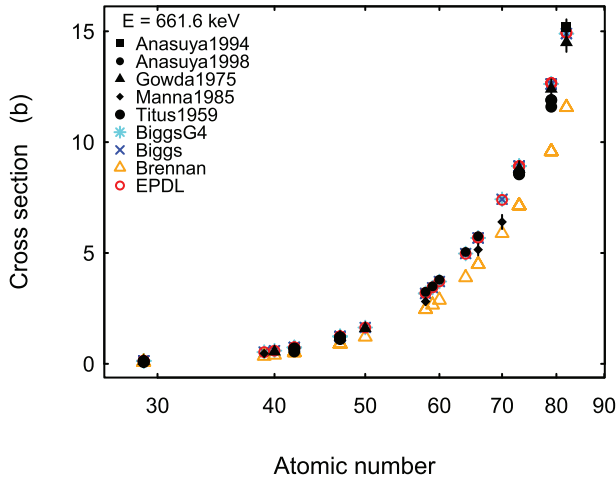


Fig. 24. Total photoionization cross section at 661.6 keV as a function of the atomic number  $Z$ .

TABLE IV  
EFFICIENCY CALCULATED WITH RESPECT TO EXPERIMENTAL  
OR SEMI-EMPIRICAL REFERENCE DATA

Model	Experimental			Semi-empirical		
	Pass	Fail	Efficiency	Pass	Fail	Efficiency
BiggsG4	33	9	$0.79 \pm 0.06$	171	128	$0.57 \pm 0.03$
Biggs	33	9	$0.79 \pm 0.06$	171	128	$0.57 \pm 0.03$
EPDL	36	6	$0.86 \pm 0.05$	180	119	$0.60 \pm 0.03$
Penelope	36	6	$0.86 \pm 0.05$	180	119	$0.60 \pm 0.03$
PHOTX	34	8	$0.81 \pm 0.06$	176	122	$0.59 \pm 0.03$
Scotfield	34	8	$0.81 \pm 0.06$	173	120	$0.59 \pm 0.03$
Storm	30	12	$0.71 \pm 0.07$	175	123	$0.59 \pm 0.03$
XCOM	35	7	$0.83 \pm 0.06$	178	120	$0.60 \pm 0.03$

TABLE V  
TEST OF EQUIVALENT COMPATIBILITY OF CALCULATED TOTAL CROSS  
SECTIONS WITH EXPERIMENTAL OR SEMI-EMPIRICAL DATA

Model	Test					
	Fisher	$\chi^2$	Boschloo	Z-pooled	Santner	Barnard
Biggs	0.0110	0.0081	0.0099	0.0090	0.0090	0.0072
EPDL	0.0010	0.0013	0.0018	0.0023	0.0018	0.0008
PHOTX	0.0064	0.0066	0.0063	0.0075	0.0080	0.0056
Scotfield	0.0063	0.0063	0.0061	0.0073	0.0079	0.0053
Storm	0.1319	0.1189	0.1219	0.1223	0.1278	0.1247
XCOM	0.0034	0.0032	0.0037	0.0041	0.0042	0.0024

The efficiency at reproducing experimental and semi-empirical reference data is reported in Table IV for all compilations covering the selected energy range. One observes that it is systematically lower, when semi-empirical data are considered as a reference in the comparison; the Wald-Wolfowitz test rejects the hypothesis of randomness of the sequence of results associated with experimental and semi-empirical references with 0.01 significance.

Categorical analysis performed over the compatibility of cross section calculations with experimental and semi-empirical reference data confirms that the observed difference is statistically significant in all cases, with the exception of the Storm and Israel compilation. The p-values resulting from different tests over contingency tables are listed in Table V. The null hypothesis of equivalent compatibility with reference data

TABLE VI  
P-VALUES FROM DIFFERENT TESTS COMPARING THE COMPATIBILITY  
WITH EXPERIMENT OF TOTAL CROSS SECTION MODELS WITH EXTENDED  
ENERGY COVERAGE AND EPDL, FOR ENERGIES ABOVE 1 KEV

Model	Test					
	Fisher	$\chi^2$	Boschloo	Z-pooled	Santner	Barnard
BiggsG4	0.570	0.393	0.427	0.427	0.586	0.942
Biggs	0.570	0.393	0.427	0.427	0.586	0.942
Penelope	1	1	1	1	1	1
PHOTX	0.771	0.558	0.608	0.639	0.744	0.997
Scotfield	0.771	0.558	0.608	0.639	0.744	0.997
Storm	0.183	0.111	0.130	0.122	0.230	0.138
XCOM	1	0.763	1	0.850	0.913	1

is rejected by all tests with 0.01 significance in the comparison involving cross sections based on Scotfield's 1973 calculations (EPDL, PHOTX, XCOM and Scotfield's own tabulations). For the comparison concerning Biggs-Lighthill cross sections, the null hypothesis is rejected by all unconditional tests and by Pearson's  $\chi^2$  tests, while it is not rejected by Fisher's exact test, which is known to be more conservative than unconditional tests. The insensitivity of the Storm and Israel model to the type of reference data to which it is compared is related to its overall lower compatibility with experiment reported in Table IV.

From these results one can infer that the use of semi-empirical data as a reference in the comparison with photoelectric cross sections would introduce systematic effects in the validation process.

All the analyses reported in the following sections concern experimental data samples only.

### B. Evaluation of Total Cross Section Compilations with Wide Energy Coverage

Some of the total cross section models considered in this study cover a wide energy range: those based on Scotfield's 1973 non-relativistic calculations (including EPDL, PHOTX and XCOM compilations), Storm and Israel's compilation and Biggs-Lighthill's parameterization, both in its original form and in the modified version used by Geant4. Their extended applicability has contributed to their extensive use in particle transport codes.

General purpose Monte Carlo codes have traditionally handled photon interactions above 1 keV; extensions to lower energies have been included only relatively recently in some of them. The validation process has investigated the ability of these cross section compilations to reproduce experimental data as a function of energy, with special attention devoted to characterizing the behaviour at low energies, below 1 keV.

The efficiency of total cross section models applicable from 1 keV up to the highest energy measurements included in the experimental sample (approximately 1.2 MeV) is reported in Table IV. The largest efficiency is achieved by EPDL (which is also the basis of Penelope's tabulations). Categorical tests based on contingency tables, summarized in Table VI, show that the differences in compatibility with experiment between the various models and EPDL are not statistically significant in this energy range.

TABLE VII  
EFFICIENCY BELOW 1 keV OF TOTAL CROSS SECTION MODELS WITH  
EXTENDED COVERAGE

$E_{min}$	$E_{max}$	Model	Pass	Fail	Efficiency
10 eV	1 keV	BiggsG4	12	31	$0.28 \pm 0.07$
		Biggs	13	30	$0.30 \pm 0.07$
		EPDL	18	25	$0.42 \pm 0.07$
100 eV	1 keV	BiggsG4	14	20	$0.41 \pm 0.08$
		Biggs	22	12	$0.65 \pm 0.08$
		EPDL	23	11	$0.68 \pm 0.08$
150 eV	1 keV	BiggsG4	14	12	$0.54 \pm 0.09$
		Biggs	17	9	$0.65 \pm 0.09$
		EPDL	19	7	$0.73 \pm 0.08$
250 eV	1 keV	BiggsG4	14	7	$0.67 \pm 0.10$
		Biggs	14	7	$0.67 \pm 0.10$
		EPDL	15	6	$0.71 \pm 0.09$

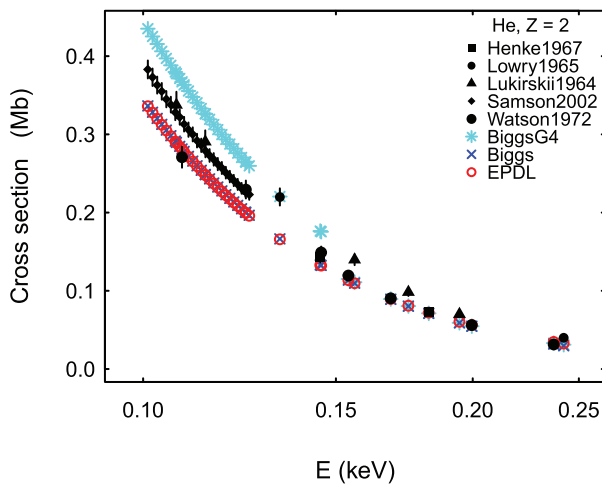


Fig. 25. Total photoionization cross section for helium as a function of photon energy, above 100 eV: original and modified Biggs-Lighthill parameterizations exhibit different behavior with respect to experimental data.

One observes some small differences in Table IV regarding the efficiencies of cross section compilations derived from Scofield's 1973 calculations. They are due to differences in the granularity of the energy grid at which cross sections are tabulated, which affects the precision of interpolation.

Only EPDL and Biggs-Lighthill's parameterization cover the whole energy range corresponding to the experimental data sample, including energies below 1 keV; their efficiencies are reported in Table VII for a series of low energy intervals. All models exhibit low efficiencies below approximately 100 eV; above this energy the efficiencies of cross sections based on EPDL and on the original Biggs-Lighthill's parameterization appear quite stable (compatible with statistical uncertainties) and similar, although EPDL ones are always larger.

The modified coefficients of Biggs and Lighthill's parameterization implemented in Geant4 do not appear to improve the compatibility with experiment of the calculated cross sections; discrepancies with respect to experimental data are qualitatively visible in Figs. 1 and 25–28. Cross sections calculated with the original coefficients appear unable to reproduce experimental data consistently in the very low energy range, below a few tens of eV: a few examples are shown in Figs. 29–32.

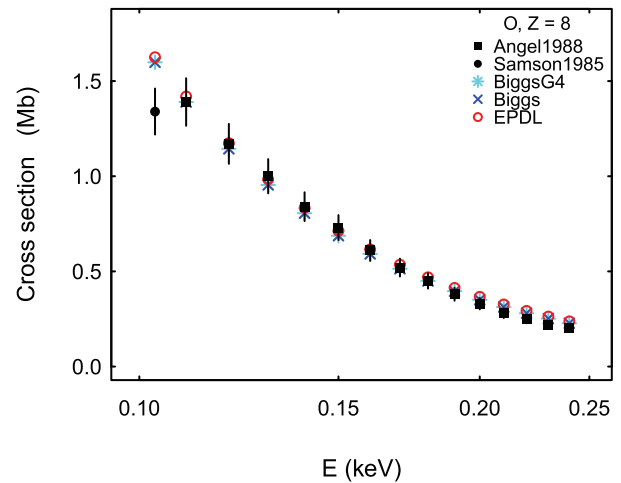


Fig. 26. Total photoionization cross section for oxygen as a function of photon energy, above 100 eV: original and modified Biggs-Lighthill parameterizations exhibit different behavior with respect to experimental data.

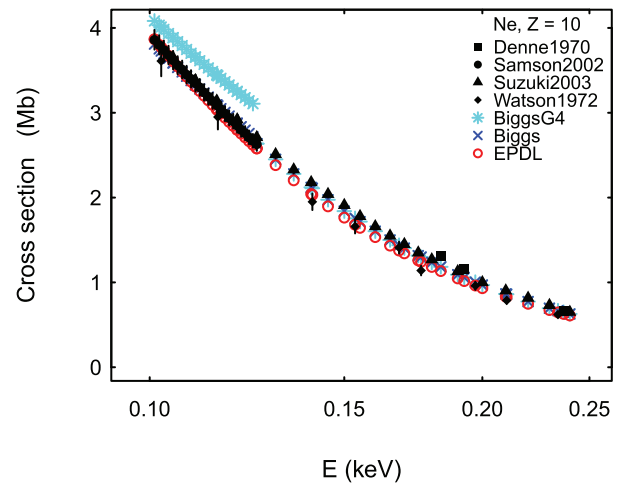


Fig. 27. Total photoionization cross section for neon as a function of photon energy, above 100 eV: original and modified Biggs-Lighthill parameterizations exhibit different behavior with respect to experimental data.

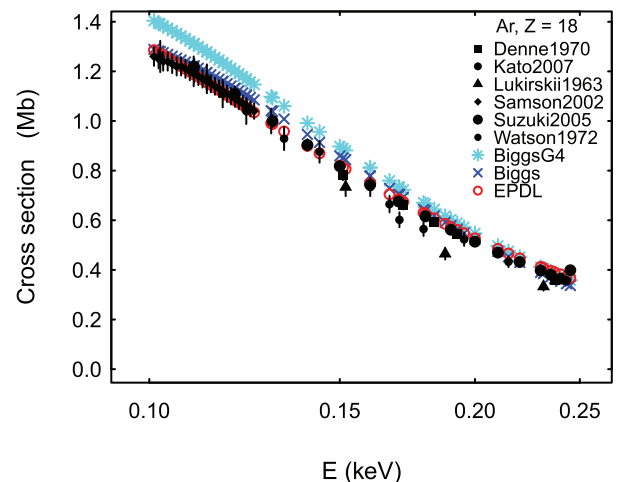


Fig. 28. Total photoionization cross section for argon as a function of photon energy, above 100 eV: original and modified Biggs-Lighthill parameterizations exhibit different behavior with respect to experimental data.

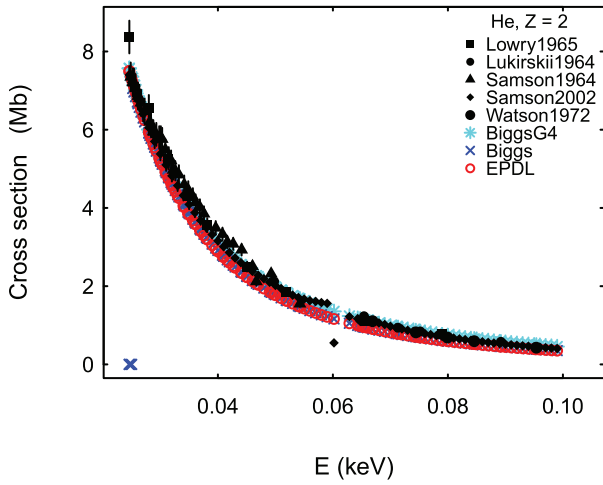


Fig. 29. Total photoionization cross section for helium as a function of photon energy, below 100 eV: original and modified Biggs-Lighthill parameterizations exhibit different behavior with respect to experimental data.

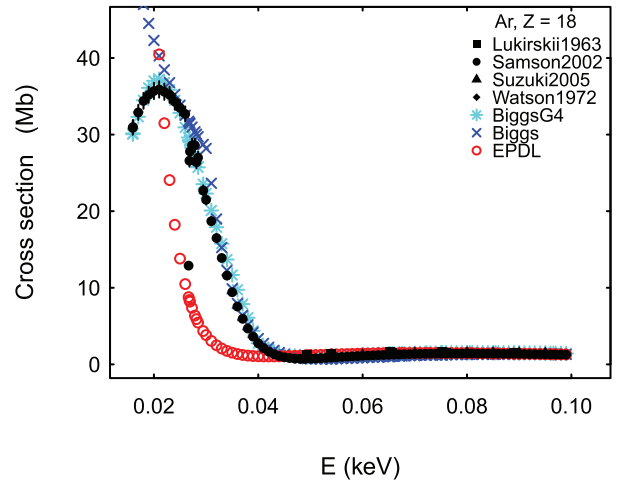


Fig. 32. Total photoionization cross section for argon as a function of photon energy, below 100 eV: original and modified Biggs-Lighthill parameterizations exhibit different behavior with respect to experimental data.

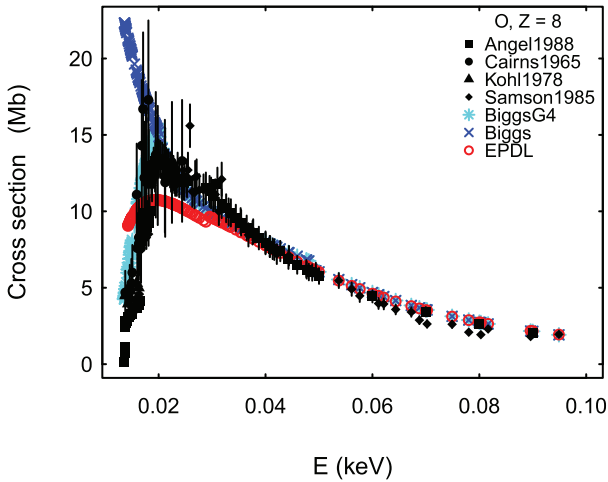


Fig. 30. Total photoionization cross section for oxygen as a function of photon energy, below 100 eV: original and modified Biggs-Lighthill parameterizations exhibit different behavior with respect to experimental data.

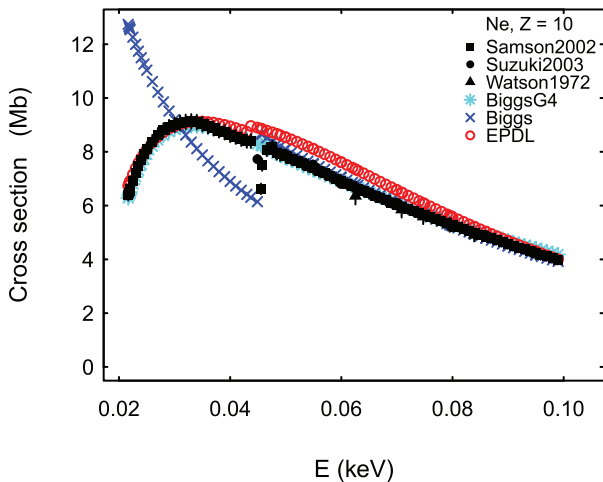


Fig. 31. Total photoionization cross section for neon as a function of photon energy, below 100 eV: original and modified Biggs-Lighthill parameterizations exhibit different behavior with respect to experimental data.

TABLE VIII  
EFFICIENCY OF TOTAL CROSS SECTIONS BASED ON ORIGINAL AND MODIFIED BIGGS-LIGHTHILL PARAMETERIZATION, LIMITED TO THE TEST CASES WHERE THEY DIFFER

Model	E < 100 eV			E ≥ 100 eV		
	Pass	Fail	Efficiency	Pass	Fail	Efficiency
BiggsG4	6	16	0.27 ± 0.09	1	13	0.07 ± 0.07
Biggs	2	20	0.09 ± 0.07	10	4	0.71 ± 0.12
EPDL	5	17	0.23 ± 0.09	11	3	0.79 ± 0.11
RTAB	7	15	0.32 ± 0.10	11	3	0.79 ± 0.11

TABLE IX  
COMPARISON OF THE COMPATIBILITY WITH EXPERIMENT OF TOTAL CROSS SECTIONS BASED ON ORIGINAL AND MODIFIED BIGGS-LIGHTHILL PARAMETERIZATION

Test	p-value, E < 100 eV	p-value, E ≥ 100 eV
Fisher	0.240	0.0013
Z-pooled	0.137	0.0004
Boschloo	0.174	0.0004
Santner	0.259	0.0004
Barnard	0.223	0.0004

A dedicated statistical analysis was performed to quantify whether the difference in compatibility with experiment between the two parameterizations is significant. For this purpose, the data sample was limited to test cases where the two cross section calculations produce different values at the same energy of an experimental measurement; these test cases concern only noble gases, oxygen and hydrogen. Two test cases were considered: one involving energies below 100 eV and one concerning measurements above or equal to 100 eV. The corresponding efficiencies are reported in Table VIII, along with EPDL and RTAB results: the original Biggs and Lighthill parameterization, as well as EPDL and RTAB compilations, exhibit low efficiencies in the lower energy range, while the efficiency of EPDL, RTAB and the original Biggs and Lighthill parameterization achieve substantially better compatibility with experiment above 100 eV. It is worth recalling that this test



TABLE X  
P-VALUES CONCERNING THE COMPARISON OF THE COMPATIBILITY WITH EXPERIMENT OF SPECIFIC TOTAL CROSS SECTION MODELS AND EPDL

Model	Energy range		p-values from tests over contingency tables					
	$E_{min}$	$E_{max}$	Fisher	$\chi^2$	Boschloo	Z-pooled	Santner	Barnard
Brennan	30 eV	700 keV	0.006	0.004	0.005	0.004	0.004	0.004
	150 eV	700 keV	0.004	0.002	0.003	0.003	0.003	0.002
Chantler	10 eV	433 keV	0.0097	0.006	0.008	0.008	0.007	0.00999
	150 eV	433 keV	0.006	0.003	0.005	0.005	0.006	0.003
Ebel	1 keV	300 keV	1	0.739	0.836	0.834	0.897	0.994
Elam	100 eV	1 MeV	0.124	0.084	0.093	0.093	0.133	0.087
Henke	10 eV	30 keV	0.151	0.101	0.109	0.110	0.128	0.108
	150 eV	30 keV	0.049	0.027	0.031	0.031	0.032	0.030
McMaster	1 keV	700 keV	0.041	0.033	0.033	0.040	0.071	0.039
	10 eV	300 keV	1	0.843	0.828	0.917	0.920	0.997
RTAB	150 eV	300 keV	1	0.793	1	0.880	0.903	0.870
	1 keV	100 MeV	0.183	0.111	0.130	0.122	0.230	0.138
Storm-Israel	100 eV	1 MeV	1	0.843	0.828	0.917	0.920	0.997
XCOM-DB	1 keV	500 keV	0.371	0.232	0.273	0.273	0.403	0.371

concerns data below 1 keV, since the coefficients of the two parameterizations differ only in the low energy range.

The results of the analysis of these categorical data, summarized in Table IX, show that above 100 eV the null hypothesis of equivalent compatibility with experiment of the two parameterizations is rejected with 0.01 significance by all the tests applied to the associated contingency table, while it is not rejected below 100 eV.

These results suggest reverting to the original coefficients of Biggs-Lighthill's parameterization in Geant4 for improved accuracy of the physics models that use cross sections based on it, at least at energies above a few tens of eV.

### C. Evaluation of Specific Total Cross Section Compilations

This investigation addresses the issue of whether specialized calculation models produce more accurate results than extensive cross section compilations in the limited energy range they cover and therefore deserve to partially replace the extensive cross section compilations currently used in general purpose Monte Carlo codes.

The total cross section compilations listed in Table IV cover different energy ranges, therefore the evaluation of their intrinsic capabilities requires specific validation tests limited to the energy range of applicability of each of them.

Efficiencies pertinent to each cross section model are reported in detail in the following subsections; they are calculated over the subset of the experimental data sample consistent with the applicability of each model. The associated tables show the model especially addressed in each test in *italic*.

The compatibility with experiment of each specialized cross section model has been compared with that of EPDL, which is the compilation with the highest efficiency in Table IV, limited to the energy range where the model subject to evaluation is applicable. The results of the statistical analysis that addresses this issue are summarized in Table X; they are discussed in detail in the following subsections.

1) *Brennan and Cowan's Cross Sections*: these cross sections cover the energy range between 30 eV and 700 keV, and are limited to atomic number greater than 2. The efficiencies listed in Table XI show lower capability of Brennan

TABLE XI  
TOTAL CROSS SECTIONS: EFFICIENCY BETWEEN 1 KEV AND 300 KEV

Model	$E_{min} = 10$ eV			$E_{min} = 150$ eV		
	Pass	Fail	Efficiency	Pass	Fail	Efficiency
BiggsG4	38	23	$0.62 \pm 0.06$	37	16	$0.70 \pm 0.06$
Biggs	38	23	$0.62 \pm 0.06$	40	13	$0.75 \pm 0.06$
<i>Brennan</i>	26	35	$0.43 \pm 0.06$	27	26	$0.51 \pm 0.07$
EPDL	42	19	$0.69 \pm 0.06$	42	11	$0.79 \pm 0.06$

TABLE XII  
TOTAL CROSS SECTIONS: EFFICIENCY BETWEEN 100 EV AND 1 MEV

Model	$E_{min} = 10$ eV			$E_{min} = 150$ eV		
	Pass	Fail	Efficiency	Pass	Fail	Efficiency
BiggsG4	34	24	$0.50 \pm 0.06$	36	15	$0.71 \pm 0.06$
Biggs	35	33	$0.51 \pm 0.06$	39	12	$0.76 \pm 0.06$
<i>Chantler</i>	24	44	$0.35 \pm 0.06$	27	24	$0.53 \pm 0.07$
EPDL	40	28	$0.59 \pm 0.06$	41	10	$0.80 \pm 0.06$

and Cowan's cross sections with respect to other compilations applicable to this energy range.

According to the results of the categorical analysis listed in Table X, the compatibility with experiment of these cross sections and EPDL ones is rejected with 0.01 significance, both over the entire energy interval of applicability and excluding test cases in the lower energy end.

From this analysis one can infer that Brennan and Cowan's calculations do not achieve greater accuracy than Scofield's 1973 non-relativistic ones tabulated in EPDL.

2) *Chantler's Total Cross Sections*: these calculations cover energies between 10 eV and 433 keV. The efficiency of the cross section models applicable in this energy range is reported in Table XII: it is lower for Chantler's cross sections than for other options applicable to this energy interval. The p-values listed in Table X show that the hypothesis of equivalent compatibility with experiment of Chantler's cross sections with respect to EPDL ones is rejected by all tests, both over the entire energy interval and excluding the lower energy end.

3) *Ebel's Total Cross Sections*: this parameterization is applicable to energies between 1 keV and 300 keV. The efficiencies listed in Table XIII, which are calculated for all

TABLE XIII

TOTAL CROSS SECTIONS: EFFICIENCY BETWEEN 1 KEV AND 300 KEV

Model	Pass	Fail	Efficiency
BiggsG4	22	8	$0.73 \pm 0.08$
Biggs	22	8	$0.73 \pm 0.08$
Brennan	19	11	$0.63 \pm 0.08$
Chantler	17	13	$0.57 \pm 0.09$
<i>Ebel</i>	24	6	$0.80 \pm 0.07$
Elam	23	7	$0.77 \pm 0.08$
EPDL	25	5	$0.83 \pm 0.07$
PHOTX	23	7	$0.77 \pm 0.08$
Scotfield	23	7	$0.77 \pm 0.08$
Storm	19	11	$0.63 \pm 0.08$
Veigele	17	13	$0.57 \pm 0.09$
<i>XCOM-DB</i>	21	9	$0.70 \pm 0.08$
XCOM	24	6	$0.80 \pm 0.08$

TABLE XIV

TOTAL CROSS SECTIONS: EFFICIENCY BETWEEN 100 eV AND 1 MeV

Model	Pass	Fail	Efficiency
BiggsG4	40	24	$0.63 \pm 0.06$
Biggs	48	16	$0.75 \pm 0.05$
<i>Elam</i>	40	24	$0.63 \pm 0.06$
EPDL	49	15	$0.77 \pm 0.05$
<i>Veigele</i>	30	34	$0.47 \pm 0.06$

TABLE XV

TOTAL CROSS SECTIONS: EFFICIENCY BETWEEN 10/150 eV AND 30 KEV

Model	$E_{min} = 10 \text{ eV}$			$E_{min} = 150 \text{ eV}$		
	Pass	Fail	Efficiency	Pass	Fail	Efficiency
BiggsG4	18	33	$0.35 \pm 0.07$	20	14	$0.59 \pm 0.08$
Biggs	19	32	$0.37 \pm 0.07$	23	11	$0.68 \pm 0.08$
Chantler	7	44	$0.14 \pm 0.05$	10	24	$0.29 \pm 0.08$
EPDL	23	28	$0.45 \pm 0.07$	24	10	$0.71 \pm 0.08$
<i>Henke</i>	15	36	$0.29 \pm 0.06$	15	19	$0.44 \pm 0.08$
RTAB	24	27	$0.47 \pm 0.07$	23	11	$0.68 \pm 0.08$

compilations covering this energy range, show similar compatibility with experiment for Ebel's and EPDL cross sections, which is confirmed by the statistical comparison of the two categories in Table X.

4) *Elam's and Veigele's Cross Sections*: they cover energies between 100 eV and 1 MeV. The efficiencies of the models applicable in this energy range are listed in Table XIV. The results of the categorical analysis in Table X show that the hypothesis of equivalent compatibility with experiment as for EPDL cross sections is rejected for Veigele's cross sections, while it is not rejected for Elam's.

5) *Henke's Cross Sections*: this model covers energies between 10 eV and 30 keV. The efficiencies for the total cross section models applicable to these energies are listed in Table XV, which also reports values for photon energies above 150 eV. Henke's efficiency is lower; nevertheless the tests summarized in Table X do not reject the hypothesis of equivalent compatibility with experiment for Henke's and EPDL cross sections with 0.01 significance.

It is worth remarking that all models appear inadequate at reproducing experimental measurements in the lower energy

TABLE XVI

TOTAL CROSS SECTIONS: EFFICIENCY BETWEEN 10/150 eV AND 300 KEV

Model	$E_{min} = 10 \text{ eV}$			$E_{min} = 150 \text{ eV}$		
	Pass	Fail	Efficiency	Pass	Fail	Efficiency
BiggsG4	29	34	$0.46 \pm 0.06$	31	15	$0.67 \pm 0.07$
Biggs	30	33	$0.48 \pm 0.06$	34	12	$0.74 \pm 0.06$
Chantler	19	44	$0.30 \pm 0.06$	22	24	$0.48 \pm 0.07$
EPDL	35	28	$0.56 \pm 0.06$	36	10	$0.78 \pm 0.06$
Penelope	37	26	$0.59 \pm 0.06$	36	10	$0.78 \pm 0.06$
RTAB	36	27	$0.57 \pm 0.06$	35	11	$0.76 \pm 0.06$

TABLE XVII

TOTAL CROSS SECTIONS: EFFICIENCY BETWEEN 1 KEV AND 500 KEV

Model	Pass	Fail	Efficiency
BiggsG4	27	8	$0.77 \pm 0.07$
Biggs	27	8	$0.77 \pm 0.07$
Brennan	22	13	$0.63 \pm 0.08$
Elam	28	7	$0.80 \pm 0.07$
EPDL	30	5	$0.86 \pm 0.06$
PHOTX	28	7	$0.80 \pm 0.07$
Scotfield	28	7	$0.80 \pm 0.07$
Storm	24	11	$0.69 \pm 0.08$
Veigele	22	13	$0.63 \pm 0.08$
<i>XCOM-DB</i>	26	9	$0.74 \pm 0.07$
XCOM	29	6	$0.83 \pm 0.06$

range, while they perform better above 150 eV. Chantler's calculations remain largely incompatible with experiment even in the higher energy interval.

6) *RTAB Total Cross Sections*: this compilation covers energies between approximately 10 eV and 300 keV. The efficiencies for the total cross section models applicable to these energies are listed in Table XVI, which also reports values for photon energies above 150 eV. The efficiency for RTAB cross sections is similar to that obtained with EPDL; the results of the statistical analysis of the related contingency table are consistent with this observation.

It is worth remarking that all models appear inadequate at reproducing experimental measurements in the lower energy range, while they perform better above 150 eV. Chantler's calculations remain largely incompatible with experiment even in the higher energy interval.

7) *XCOM-DB Cross Sections*: this compilation, encompassed in the DABAX database, concerns energies between 1 keV and 500 keV. The efficiencies of the applicable total cross section models are listed in Table XVII. Neither the tests on contingency tables concerning XCOM-DB and EPDL (reported in Table X), nor those concerning XCOM-DB and the standard XCOM tabulation reject the hypothesis of equivalent compatibility with experiment.

The different efficiency for cross sections based on XCOM-DB and standard XCOM tabulations could derive from a different energy grid of the two tabulations, which affects the cross section values calculated by interpolation, or from a different version of XCOM used as a basis for XCOM-DB. The XCOM version used for creating the XCOM-DB tabulations could not be retrieved in the DABAX documentation.

The quantitative validation analysis documented here shows that none of the specialized cross section models provides better

TABLE XVIII  
EFFICIENCY OF TOTAL CROSS SECTION MODELS IN THE LOW ENERGY RANGE

Model	10-100 eV			10-150 eV		
	Pass	Fail	Efficiency	Pass	Fail	Efficiency
BiggsG4	9	22	0.29 ± 0.08	7	28	0.20 ± 0.07
Biggs	5	26	0.16 ± 0.07	8	27	0.23 ± 0.07
Chantler	4	27	0.13 ± 0.08	5	30	0.14 ± 0.06
EPDL	8	23	0.26 ± 0.06	10	25	0.29 ± 0.07
Henke	9	22	0.29 ± 0.08	11	24	0.31 ± 0.08
Penelope	10	21	0.32 ± 0.08	12	23	0.34 ± 0.08
RTAB	10	21	0.32 ± 0.08	12	23	0.34 ± 0.08

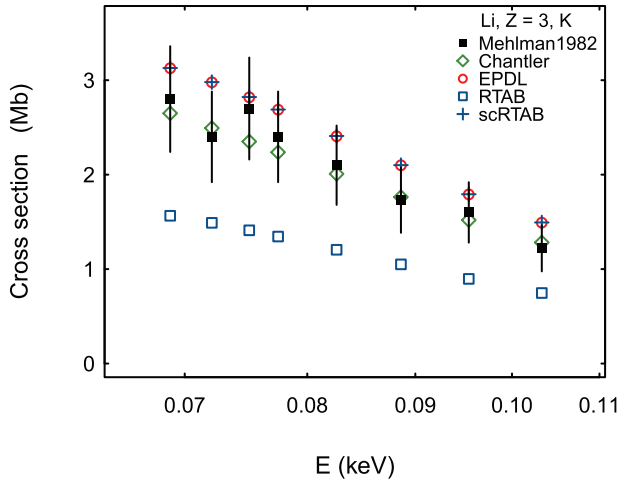


Fig. 33. Cross section for the photoionization of the K shell of lithium as a function of photon energy.

accuracy than EPDL tabulations: some of them exhibit significantly worse compatibility with experiment, while others are at most statistically equivalent to EPDL at reproducing experimental measurements. Therefore their use in their specific range of applicability at the place of EPDL would not improve the accuracy of general purpose Monte Carlo codes.

#### D. Evaluation of Total Cross Sections at Low Energy

The results documented in the previous sections hint that neither the models applicable to an extended energy range nor those covering specific energies appear capable of reproducing cross sections measurements at low energies.

$\chi^2$  tests performed over a data sample with energies between 10 eV and 150 eV quantify the capability of all models applicable in that energy range to reproduce experimental data. The resulting efficiencies are documented in Table XVIII; they confirm that none of the cross section compilation is suitable for accurate simulation of the photoelectric effect in this energy range.

### VIII. RESULTS OF PARTIAL CROSS SECTION VALIDATION

Figs. 33 to 51 illustrate some examples of calculated and experimental cross sections for inner and outer shell photoionization.

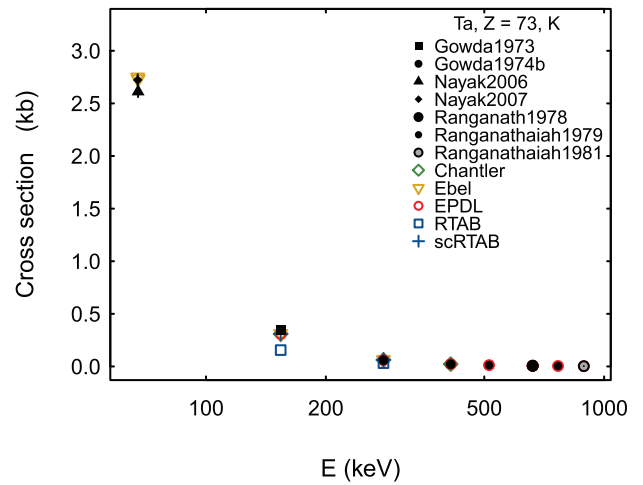


Fig. 34. Cross section for the photoionization of the K shell of tantalum as a function of photon energy.

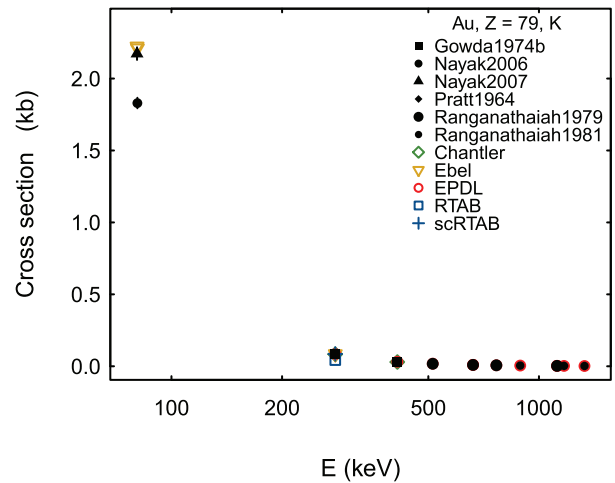


Fig. 35. Cross section for the photoionization of the K shell of gold as a function of photon energy.

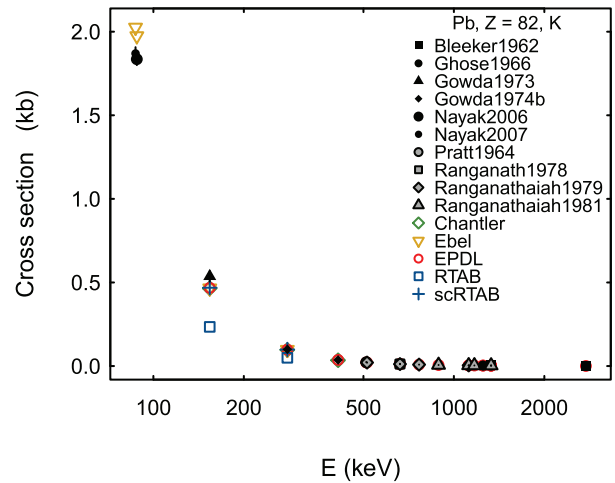


Fig. 36. Cross section for the photoionization of the K shell of lead as a function of photon energy.

A systematic discrepancy of RTAB shell cross sections with respect to experimental data is observed, which hints at a missing multiplicative factor in the tabulated values. RTAB cross

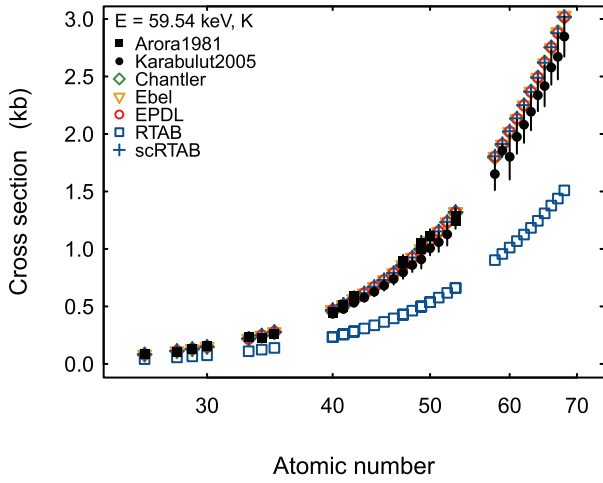


Fig. 37. Cross section for the photoionization of the K shell at 59.54 keV as a function of the atomic number.

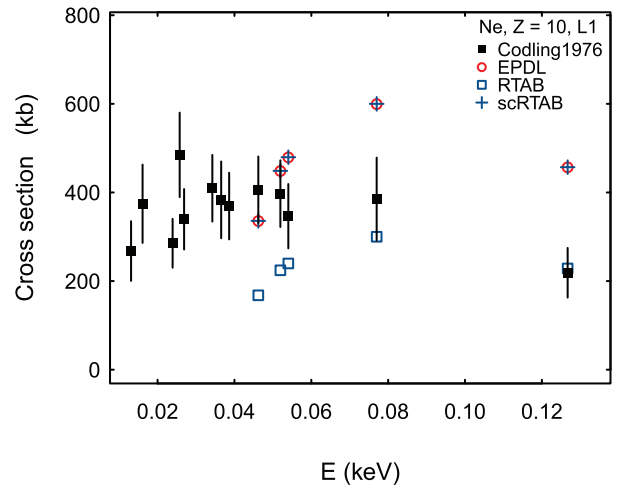


Fig. 40. Cross section for the photoionization of the  $L_1$  subshell of neon as a function of photon energy.

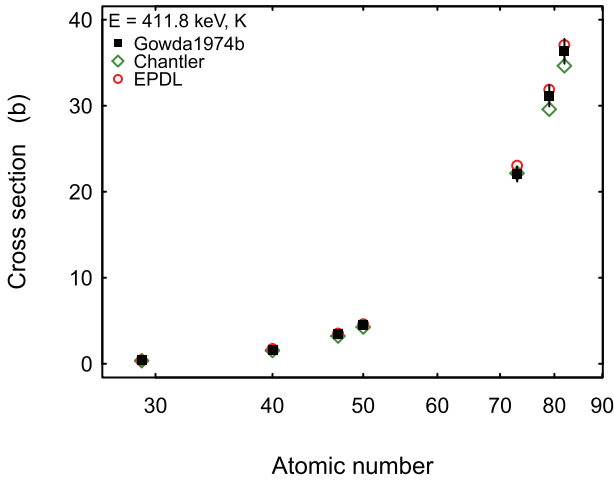


Fig. 38. Cross section for the photoionization of the K shell at 411.8 keV as a function of the atomic number.

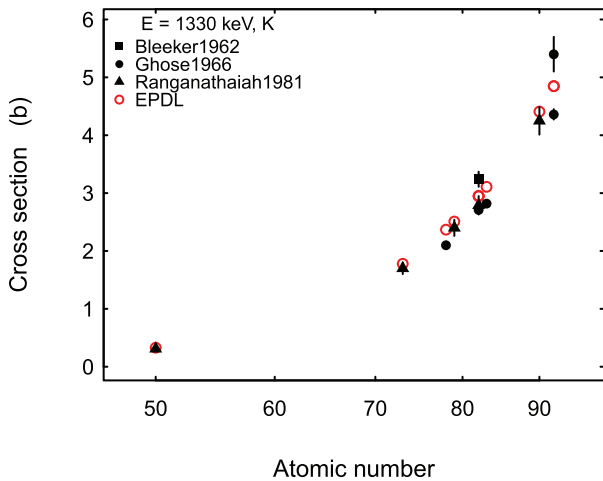


Fig. 39. Cross section for the photoionization of the K shell at 1.33 MeV as a function of the atomic number.

sections scaled by the presumed missing factor are identified in figures and in the following tables as “scRTAB”.

### A. K Shell

All but one of the experimental data sets concern energies above 1 keV.

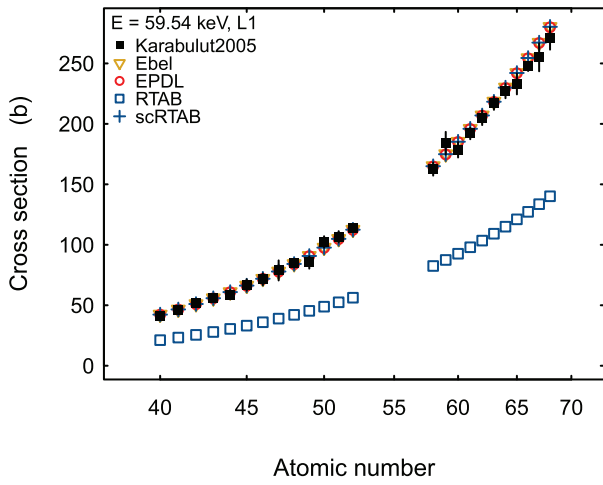
Only the compilations based on Scofield’s 1973 non-relativistic calculations cover the whole energy range of the experimental data sample above 1 keV; their efficiencies are listed in Table XIX. The lower efficiency obtained with PHOTX is related to the characteristics of the energy grid over which cross sections are tabulated.

Chantler’s, EPDL and scRTAB models are applicable to the single set of measurements below 1 keV retrieved from the literature; according to the outcome of the  $\chi^2$  test, their calculations are compatible with the experimental data set with 0.01 significance.

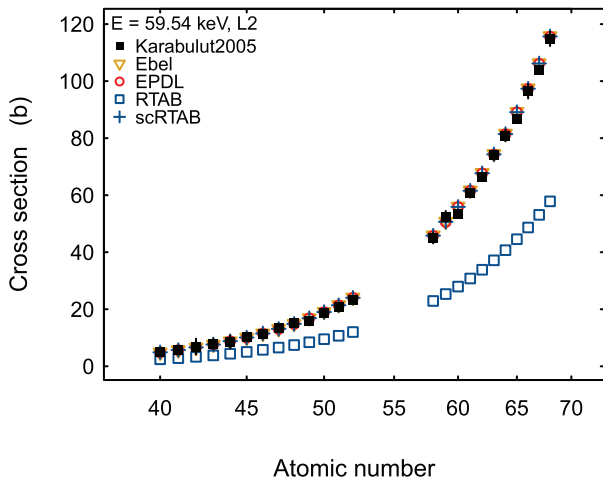
The capabilities of Chantler’s, Ebel’s and scRTAB models, which are limited to a specific energy range, are evaluated by restricting the  $\chi^2$  test of comparison with experiment to the data sample consistent with their coverage. The resulting efficiencies, calculated within 1 keV and 300 keV, which correspond to the coverage of Ebel’s model, are reported in Table XX. Since the experimental data sample includes no additional measurements between 300 keV and 433 keV, which is the highest energy tabulated by Chantler’s calculations, the results in Table XX express the outcome of the validation test for this model too.

Ebel’s parameterization appears more efficient than EPDL tabulations at reproducing experimental K shell cross sections in the energy range between 1 and 300 keV; nevertheless, the hypothesis of equivalent compatibility with experiment for these K shell cross section models is not rejected by any of the tests applied to the analysis of the associated contingency table with 0.01 significance. The failures of EPDL-based cross sections at reproducing experimental data mostly concern test cases close to absorption edges. Issues about the accuracy of related calculations of atomic binding energies

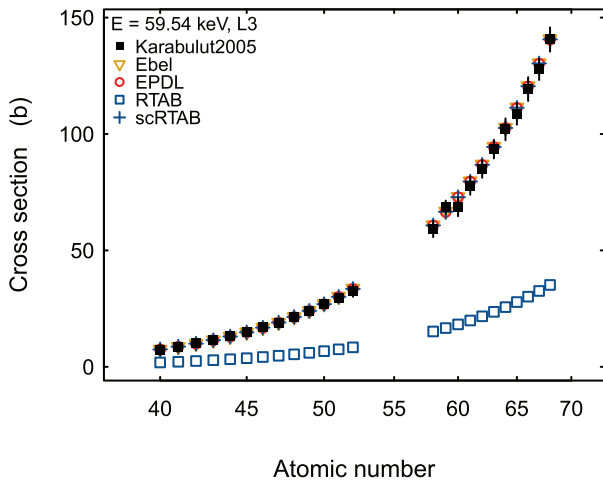




(a) L<sub>1</sub> subshell

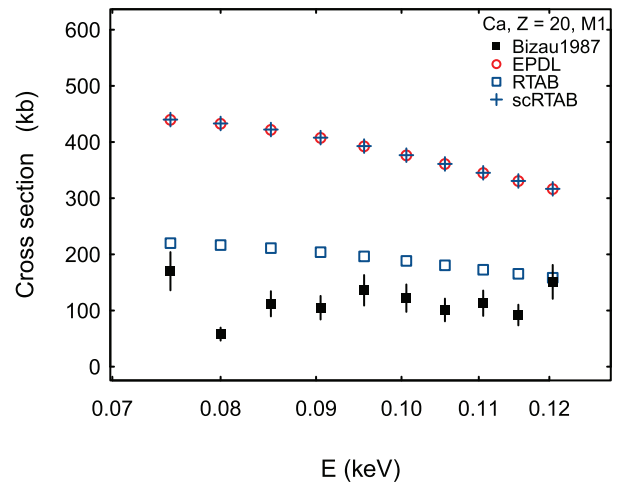


(b) L<sub>2</sub> subshell

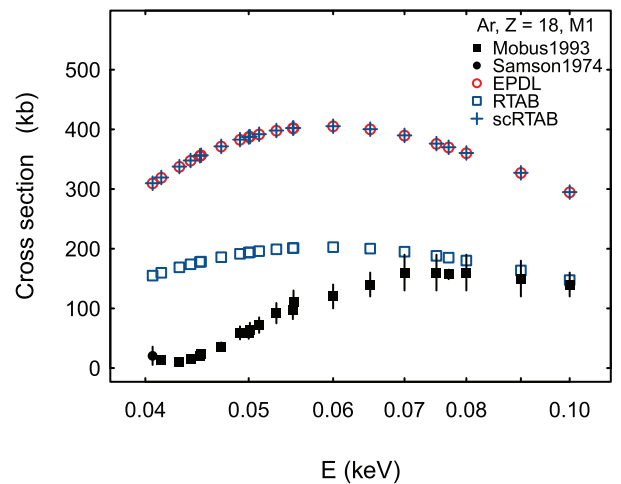


(c) L<sub>3</sub> subshell

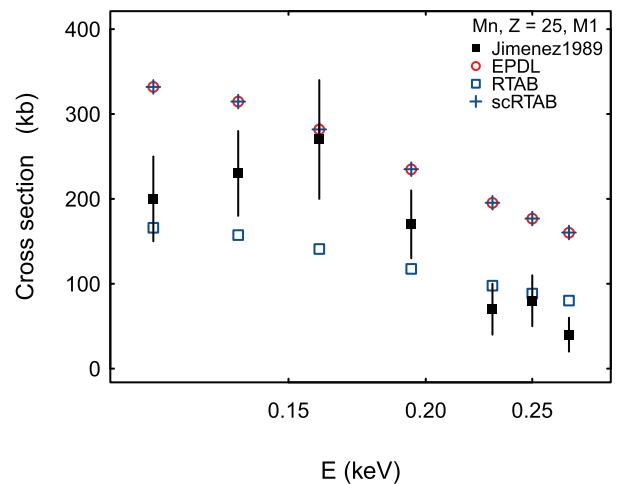
Fig. 41. Cross section for the photoionization of the L shell at 59.54 keV as a function of the atomic number. (a) L<sub>1</sub> subshell (b) L<sub>2</sub> subshell (c) L<sub>3</sub> subshell.



(a) Argon (Z=18)



(b) Calcium (Z=20)



(c) Manganese (Z=25)

Fig. 42. Cross section for the photoionization of the M<sub>1</sub> subshell of argon, calcium and manganese as a function of energy. (a) Argon (Z = 18) (b) Calcium (Z = 20) (c) Manganese (Z = 25).

collected in EADL (Evaluated Atomic Data Library) [267] were documented in [22]; as discussed in Section VI-B, experimental measurements in these regions are delicate, as

realistic estimates of experimental uncertainties are prone to be affected by systematic effects deriving from the experimental environment.

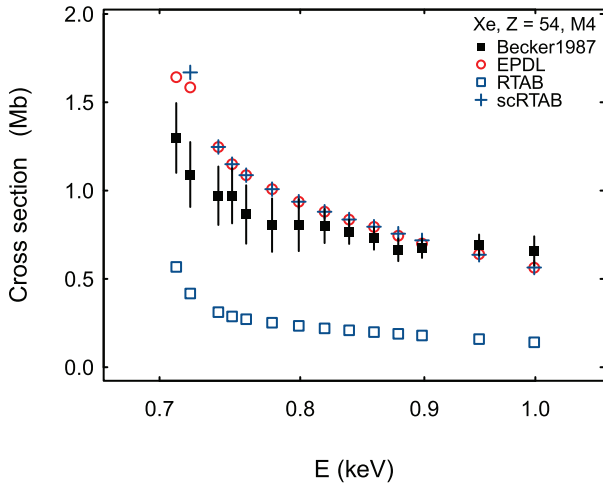


Fig. 43. Cross section for the photoionization of the M<sub>4</sub> subshell of xenon as a function of photon energy.

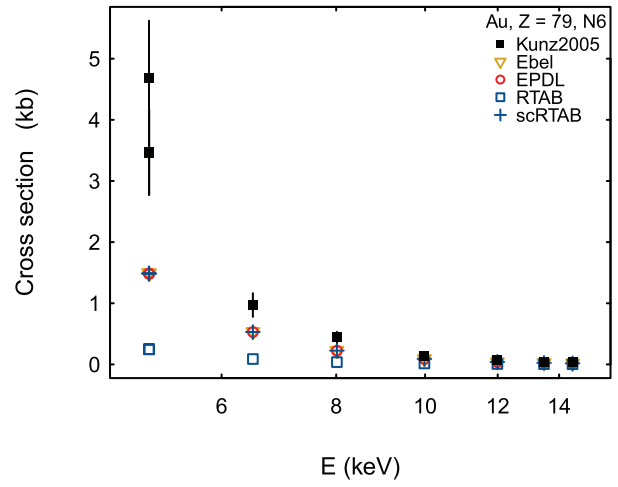


Fig. 46. Cross section for the photoionization of the N<sub>6</sub> subshell of gold as a function of photon energy.

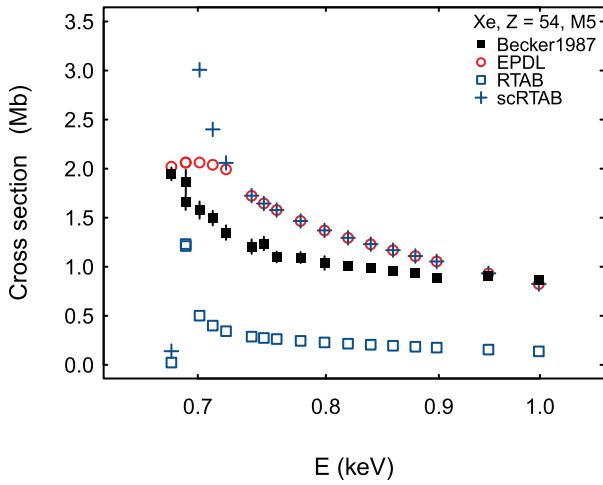


Fig. 44. Cross section for the photoionization of the M<sub>5</sub> subshell of xenon as a function of photon energy.

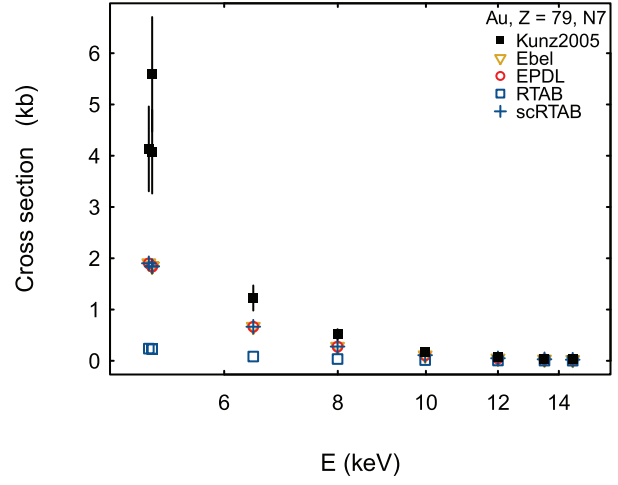


Fig. 47. Cross section for the photoionization of the N<sub>7</sub> subshell of gold as a function of photon energy.

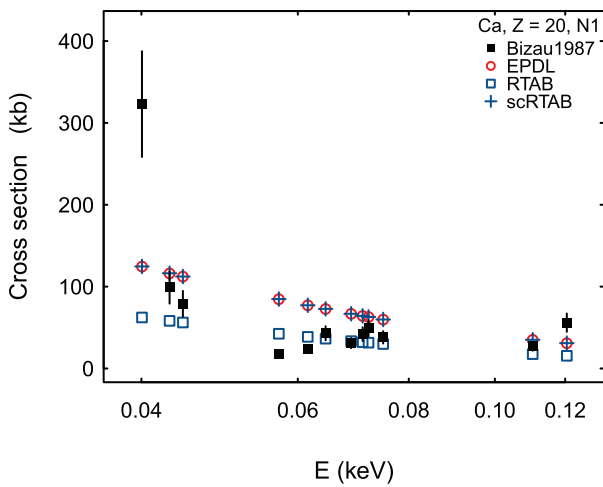


Fig. 45. Cross section for the photoionization of the N<sub>1</sub> subshell of calcium as a function of photon energy.

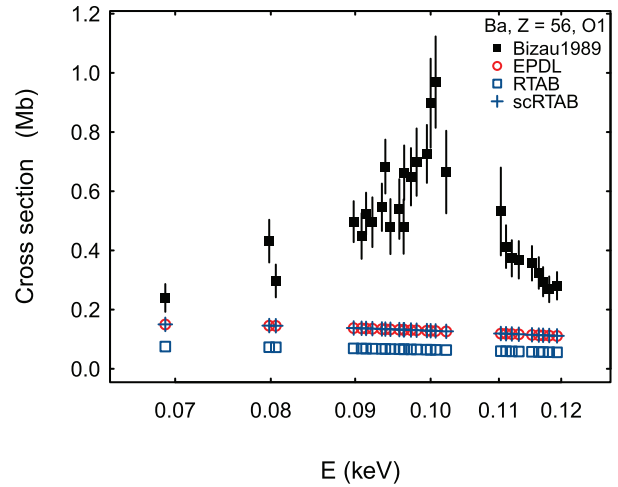


Fig. 48. Cross section for the photoionization of the O<sub>1</sub> subshell of barium as a function of photon energy.

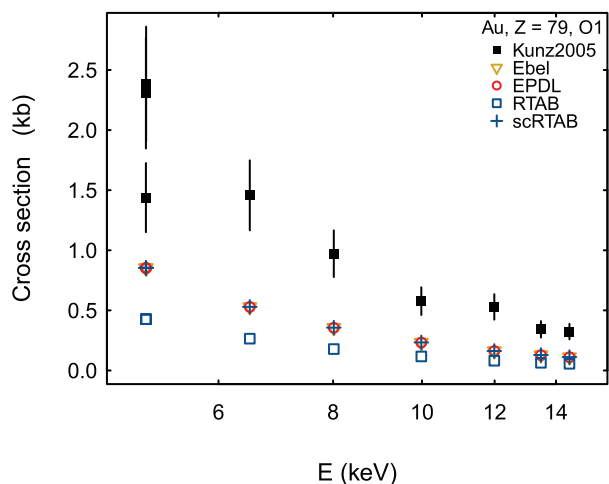


Fig. 49. Cross section for the photoionization of the O<sub>1</sub> subshell of gold as a function of photon energy.

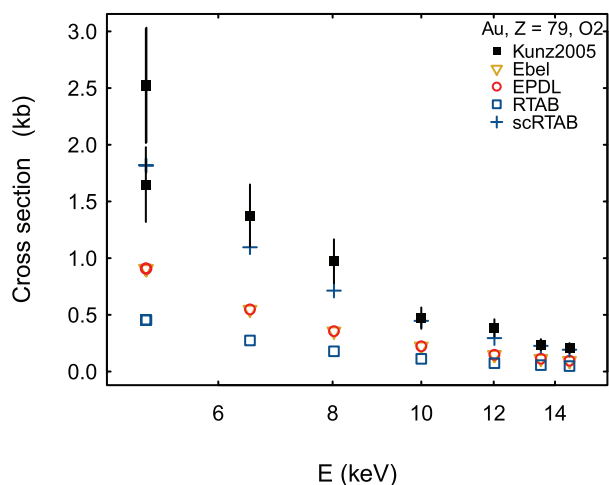


Fig. 50. Cross section for the photoionization of the O<sub>2</sub> subshell of gold as a function of photon energy.

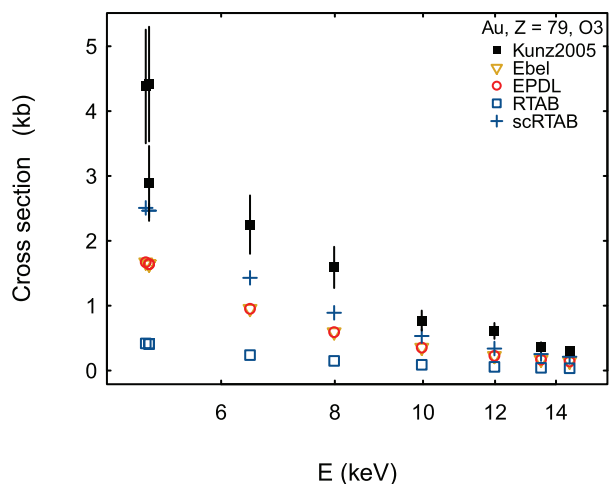


Fig. 51. Cross section for the photoionization of the O<sub>3</sub> subshell of gold as a function of photon energy.

TABLE XIX  
K SHELL CROSS SECTION EFFICIENCY OF MODELS BASED ON SCOFIELD  
1973 NON-RELATIVISTIC CALCULATIONS

Model	Pass	Fail	Efficiency
EPDL	62	17	$0.78 \pm 0.05$
Penelope	62	17	$0.78 \pm 0.05$
PHOTX	52	27	$0.66 \pm 0.05$
Scotfield	62	17	$0.78 \pm 0.05$

TABLE XX  
EFFICIENCY OF K SHELL CROSS SECTION MODELS FOR ENERGIES  
BETWEEN 1 KEV AND 300 KEV

Model	Pass	Fail	Efficiency
Chantler	45	13	$0.78 \pm 0.05$
Ebel	50	8	$0.86 \pm 0.05$
EPDL	45	13	$0.78 \pm 0.05$
Penelope	45	13	$0.78 \pm 0.05$
PHOTX	47	11	$0.81 \pm 0.05$
scRTAB	45	13	$0.78 \pm 0.05$
Scotfield	45	13	$0.78 \pm 0.05$

TABLE XXI  
EFFICIENCY OF L SHELL CROSS SECTION MODELS ABOVE 1 KEV

Subshell	Model	Pass	Fail	Efficiency
L <sub>1</sub>	Ebel	24	0	1.00 - 0.04
	EPDL	24	0	1.00 - 0.04
	Penelope	24	0	1.00 - 0.04
	PHOTX	24	0	1.00 - 0.04
	scRTAB	24	0	1.00 - 0.04
	Scotfield	24	0	1.00 - 0.04
L <sub>2</sub>	Ebel	23	1	$0.96 \pm 0.04$
	EPDL	23	1	$0.96 \pm 0.04$
	Penelope	23	1	$0.96 \pm 0.04$
	PHOTX	23	1	$0.96 \pm 0.04$
	scRTAB	23	1	$0.96 \pm 0.04$
	Scotfield	23	1	$0.96 \pm 0.04$
L <sub>3</sub>	Ebel	27	0	1.00 - 0.03
	EPDL	27	0	1.00 - 0.03
	Penelope	27	0	1.00 - 0.03
	PHOTX	27	0	1.00 - 0.03
	scRTAB	27	0	1.00 - 0.03
	Scotfield	27	0	1.00 - 0.03

### B. L Shell

All L subshell measurements included in the experimental data sample concern energies between 1 keV and 300 keV, with the exception of one set of L<sub>1</sub> cross section data, which encompasses lower energy measurements.

Efficiencies above 1 keV are reported in Table XXI for all models dealing with L subshells. All cross section models are equally capable of reproducing the measurements in the experimental data sample.

The hypothesis of compatibility with the single experimental data set including measurements below 1 keV is rejected for all models subject to evaluation. Nevertheless, no conclusions regarding the capabilities of the cross section models at lower energies can be inferred from such a limited test; the failure over a single experimental data set could be due either to deficiency of the models or to systematic effects affecting the measurements, which cannot be investigated in terms of consistency with other, independent experimental measurements.

### C. Outer Shells

The cross sections for outer shells appear incompatible with experiment; the  $\chi^2$  test rejects the hypothesis of compatibility between calculated and experimental cross sections in most test cases. Nevertheless, caution should be exercised in drawing general conclusions from these results, as the experimental data sample for the validation of outer shells is small and for a given test case the available data often originate from a single experimental source. This test scenario is prone to be affected by systematic effects, which would hinder the reliability of the outcome of the statistical analysis.

Firm conclusions about the accuracy of cross section calculations for outer shells would require more extensive experimental data samples.

## IX. QUALITATIVE EVALUATION OF PHOTOELECTRON ANGULAR DISTRIBUTION

A preliminary appraisal of photoelectron angular distribution models has been performed, limited to the different options implemented in Geant4 and GEANT 3. Inconsistencies in the analytical formula of the latter appearing in the software documentation have been corrected in the implementation used for this evaluation.

Reports of direct measurements of photoelectron angular distribution are scarce in the literature; the narrow scope of the experimental data sample we could retrieve from the literature and the lack of adequate detail on the characteristics of the reported measurements prevent a meaningful statistical study for the validation of the models used in general purpose Monte Carlo codes.

Figs. 52 and 53 show qualitative comparisons of photoelectron angular distribution models implemented in Geant4 and GEANT 3 with experimental data; discrepancies are visible, both between calculations and experimental data and between the outcome of different model implementations, which suggest that further quantitative investigation would be useful. In the scenarios corresponding to the limited experimental sample retrieved from the literature, Geant4 angular distribution models exhibit similar behaviour; the corrected GEANT 3 model appears in some cases different from the others and qualitatively closer to measurements. Nevertheless, no conclusions can be drawn from such a narrow, qualitative test.

An indirect investigation of photoelectric differential cross section models could be pursued through testing the so-called ‘‘asymmetry parameter’’ [5]; due to its complexity, this study should be considered as the subject of a dedicated paper.

## X. CONCLUSION

An extensive set of cross section models for the simulation of photoionization has been quantitatively evaluated regarding their accuracy at reproducing experimental measurements. Since the validation analysis documented in this paper concerns cross section data libraries and parameterization methods, which are (and can be) used by any Monte Carlo codes, rather than specific implementations of the photoelectric effect, its

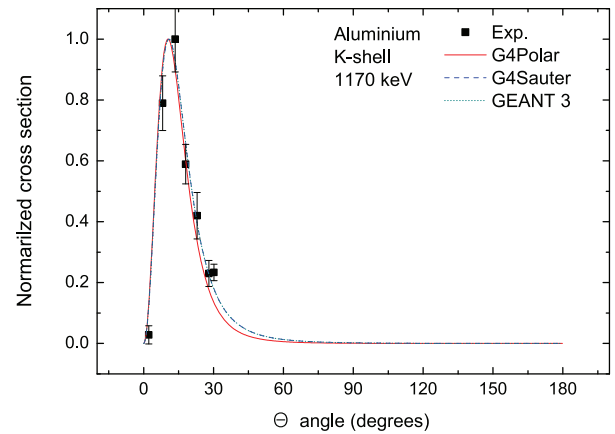


Fig. 52. Photoelectron angular distribution for aluminium, K shell, at 1.17 MeV.

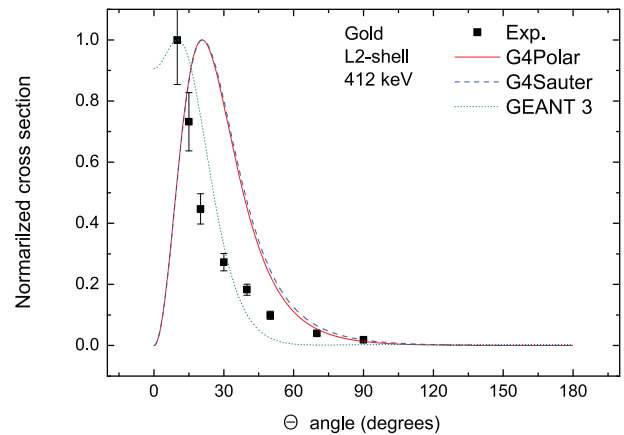


Fig. 53. Photoelectron angular distribution for gold,  $L_2$  subshell, at 412 keV.

results, and the recommendations which derive from them, are of general relevance for the simulation of the photoelectric effect.

Statistical tests against a large sample of total and partial cross section measurements demonstrate that tabulations based on Scofield’s 1973 non-relativistic calculations, which cover an extended range of energies starting from a few electronvolts, achieve the best overall capability to reproduce experiment among the models subject to test. Small differences in the results of goodness-of-fit tests are observed across different compilations (EPDL, Penelope database, PHOTX, XCOM) based on the same theoretical calculation method; they appear related to the granularity of the tabulations.

Some discrepancies with experiment of EPDL-based K shell cross sections are observed close to absorption edges, which could be related to intrinsic limitations of the calculations, also highlighted in the evaluation of related calculations of atomic binding energies; nevertheless, the possibility of systematic effects in these delicate measurements suggests caution in drawing conclusions about the reliability of EPDL in the proximity of absorption edges.



No other photoelectric cross section calculation methods, among those considered in this study, demonstrate statistically better accuracy than EPDL tabulations: neither theoretical calculations, including relativistic ones, nor empirical or theory-driven analytical parameterizations. The validation tests documented in this paper demonstrate that the compatibility with experiment of some of them is significantly worse than that of EPDL, while it is statistically equivalent for others, although restricted to the energy range they cover. Only K-shell cross sections based on Ebel's parameterization produce more accurate results than EPDL in some test cases close to absorption edges, although the difference in compatibility with experiment with respect to EPDL is not statistically significant.

Special attention has been devoted to the evaluation of total cross sections based on Biggs and Lighthill's parameterization, which is extensively used in Geant4. While the original set of parameterization coefficients produces cross sections that are statistically equivalent to EPDL regarding compatibility with experiment, the modified coefficients implemented in Geant4 significantly worsen the resulting cross section compatibility with experimental measurements above a few tens of electron-volt. The original and modified parameterization produce statistically equivalent compatibility with experiment below a few tens of electron-volt; both parameterizations, as well as other total cross section models, exhibit low efficiency in that energy range. As a result of this validation test, it is recommended that future versions of Geant4 revert to using the original coefficients of Biggs and Lighthill parameterization above a few tens of electron-volt, unless a new, more accurate parameterization is developed and quantitatively demonstrated to improve compatibility with experiment over the original one. The suggestion to implement this modification in Geant4 version 10.2, which is in preparation at the time of writing this paper, has been conveyed to the maintainers of the related Geant4 code. Given that EPDL photoelectric cross sections are already used in other Geant4 models [93], the burden of maintaining alternative cross sections based on Biggs-Lighthill's parameterization should be considered as well, as this validation test has demonstrated that even the original parameterization is not statistically superior to EPDL in reproducing experimental data.

The validation tests show that in the low energy end cross sections based on Scofield's non-relativistic calculations produce results consistent with experiment with 0.01 significance down to approximately 150 eV. Nevertheless, the limited representativity of the experimental data sample, which only includes noble gases and hydrogen, suggests caution at extrapolating this result to experimental scenarios involving solid targets.

No model, among those considered in this validation study, is able to reproduce experimental data consistently below approximately 150 eV. Their failure at low energies reflects the inadequacy of theoretical calculations in independent particle approximation and of the assumption of isolated atoms adopted in general purpose Monte Carlo codes, in conditions where effects related to the structure of the interacting material should be taken into account.

The quantitative results of the validation tests documented in this paper allow Monte Carlo code developers and experimental

users to base the selection of physics models in their simulations on sound scientific grounds.

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